

Classical Electromagnetism as a Pathway beyond Lorentzian Analogues

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- 1 EM & Gordon Metric
- 2 Beyond Lorentzian Geometry
- 3 UdW in Uniaxial Crystal
- 4 Summary

Caveats & Conventions

- Signature: $-+++$
- $G = c = \hbar = 1$
- Space-time indices: $abcd \dots$
- Boys–Post constitutive relations:

$$\begin{pmatrix} \mathbf{D} \\ \mathbf{H} \end{pmatrix} = \begin{pmatrix} \epsilon & \zeta \\ \zeta^\dagger & \mu^{-1} \end{pmatrix} \begin{pmatrix} \mathbf{E} \\ \mathbf{B} \end{pmatrix}. \quad (1)$$

- **PSA**: EM terminology *not* gravitoelectromagnetism
- **PSA**: ‘Analogue Space-Times’ \gg ‘Analogue Gravity’¹
- arXiv numbers for convenience, not to imply ‘unpublished’

¹But see, for example, Erkul & Leonhardt (2025) arXiv:[2508.11300](https://arxiv.org/abs/2508.11300), Volovik – *The Universe in a Helium Droplet*, ...

Electromagnetism and the Earliest Analogue

Reminder: Orthogonal Decomposition

For any two-form F_{ab} in four dimensional space-time and any four-velocity V^a , there exist two vector fields E^a and B^a , both orthogonal to V^a , such that

$$F_{ab} = V_a E_b - V_b E_a + \varepsilon_{abcd} V^c B^d. \quad (2)$$

- Similar statements for any possible tensor
- View our constitutive equation (1) as tensorial map of field strength $F_{ab} \xrightarrow{Z} G$, excitation tensor
- Z has 21 d.o.f.² if compatible with an action $\mathcal{L} \propto F_{ab} Z^{abcd} F_{cd}$

- $$\begin{aligned} \epsilon_V^{ab} &= -2Z^{dacb} V_d V_c, & [\mu_V^{-1}]^{ab} &= 2(*Z*)^{dacb} V_d V_c, \\ \zeta_V^{ab} &= 2(*Z)^{dacb} V_d V_c, & [\zeta_V^\dagger]^{ab} &= 2(Z*)^{dacb} V_d V_c. \end{aligned}$$

²Often the axion-part $\propto \varepsilon_{abcd}$ is omitted, but see, for example, Hehl *et al.* (2009), arXiv:[0903.1261](https://arxiv.org/abs/0903.1261).

- This directly translates to the physics of moving media!³
- An isotropic medium in its rest frame V viewed by W :

$$\begin{aligned}\epsilon_W^{bd} &= -2Z^{abcd}W_aW_c, \\ &= \mu^{-1}(g^{bd} + W^bW^d) + (\epsilon - \mu^{-1})(g^{bd}(V \cdot W)^2 \\ &\quad - (W^bV^d + V^bW^d)(V \cdot W) - V^bV^d), \\ [\mu_W^{-1}]^{bd} &= \frac{h_W^{bd}}{\mu} + (\mu^{-1} - \epsilon) \left((V \cdot W)^2 h_W^{bd} - h_W^{be} h_{ef} h_W^{fd} \right), \\ \zeta_W^{ac} &= (\epsilon - \mu^{-1})(V \cdot W) (\epsilon^{acef} W_e V_f).\end{aligned}$$

- Reminders: $(V \cdot W)^2 = \gamma_{V,W}$, $1 - 1/(\epsilon\mu) = 1 - 1/n^2$
- Isotropy is lost in the moving medium!

³A.k.a. the Fresnel–Fizeau effect; see, e.g., Post – *Formal Structure of Electromagnetics* or O'Dell - *The Electrodynamics of Magneto-Electric Media*.

The Issue with the Gordon Analogue

- In this language, the Gordon analogue, ' $g \rightarrow n$ ', would read as: Find Z such that

$$S \stackrel{!}{=} -\frac{1}{8} \int d^4x \sqrt{-\det g_{\text{eff}}} \left([g_{\text{eff}}^{-1}]^{ac} [g_{\text{eff}}^{-1}]^{bd} - [g_{\text{eff}}^{-1}]^{ad} [g_{\text{eff}}^{-1}]^{bc} \right) F_{ab} F_{cd}, \quad (3)$$

$$\text{i.e., } Z^{abcd} = \frac{1}{2} \frac{\sqrt{\det g_{\text{eff}}}}{\sqrt{\det g}} \left([g_{\text{eff}}^{-1}]^{ac} [g_{\text{eff}}^{-1}]^{bd} - [g_{\text{eff}}^{-1}]^{ad} [g_{\text{eff}}^{-1}]^{bc} \right). \quad (4)$$

- The consistency condition (for isotropic media $\epsilon = \mu$) follows from d. o. f. (g_{ab}^{eff}) \neq d. o. f. (Z^{abcd}).⁴
- **Warning!** $\nabla_{g_{\text{lab}}}$ is not metric-compatible with $\nabla_{g_{\text{eff}}}$ and *vice versa*

⁴See, for example, SeSc, Visser arXiv:[1706.06280](https://arxiv.org/abs/1706.06280) for the general case.

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- \implies **Cartographic distortions from (not just!) non-metricity**

$$q_{bca}^{\text{lab/eff}} := \nabla_a^{\text{lab/eff}} g_{bc}^{\text{eff/lab}}, \quad (5)$$

- This has often (implicitly) been re-discovered: Cummer, Fathi, Frauendiener, Thompson (arXiv:1006.3364, arXiv:1602.08341, arXiv:1705.11108); Sawicki, Trenkler, Vikman (arXiv:2412.21169); SeSc, Visser (arXiv:1808.07987)

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Issues of Coordinates

- Algebraic EM analogue of Schwarzschild:

$$\epsilon^{ab} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & \frac{1}{r(r-2M)} & 0 \\ 0 & 0 & \frac{1}{r(r-2M)\sin^2\theta} \end{pmatrix},$$

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Example: Scaled lab coordinates

Hawking temperature of a Schwarzschild BH

- Hawking temperature can be rephrased in terms of either M or ϵ , $[\mu^{-1}]$, ζ

$$\bullet T_{\text{H,eff}} = \left(\frac{\hbar}{4\pi\sqrt{\det g_{\text{lab}} \det \epsilon}} \right) \Big|_{r_{\text{eff}}} \Big|_{r_{\text{eff}}=r_+}$$

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$$\bullet T_{\text{H,lab}} = \frac{\hbar}{8\pi a^{9/2} M} = a^{-9/2} T_{\text{H,eff}}$$

- $T_{\text{H,eff}}$ conformally invariant under conformal *space-time* transformations of (M, g_{eff}) ; not under *space* transformations!

Beyond Lorentzian Geometry

Area Metrics, Finsler Metrics, and Dispersion Relations

- For QG or EM—‘area metrics’; reinterpreting Z as mapping area-forms⁵

$$\text{EM: } S_{\text{EM}} = -\frac{1}{8} \int_M d^4x \sqrt{\det G_{\text{Petrov}}^{6 \times 6}} F_{ab} Z^{abcd} F_{cd},$$

$$\text{GR+: } S_{\text{EH}} = \int_M d^4x \sqrt{\det G_{\text{Petrov}}^{6 \times 6}} 'R'.$$

\Rightarrow ‘Area metric geometry’

⁵For details: Punzi, Schuller, Witte, Wohlfarth (arXiv:[hep-th/0508170](https://arxiv.org/abs/hep-th/0508170), arXiv:[hep-th/0612141](https://arxiv.org/abs/hep-th/0612141), arXiv:[0908.1016](https://arxiv.org/abs/0908.1016)).

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- Z also tied to Finsler geometry:⁶

$$ds = F(x, y), \quad x \in M, y \in T_x M, \quad \text{homogeneous of degree 1 in } y$$

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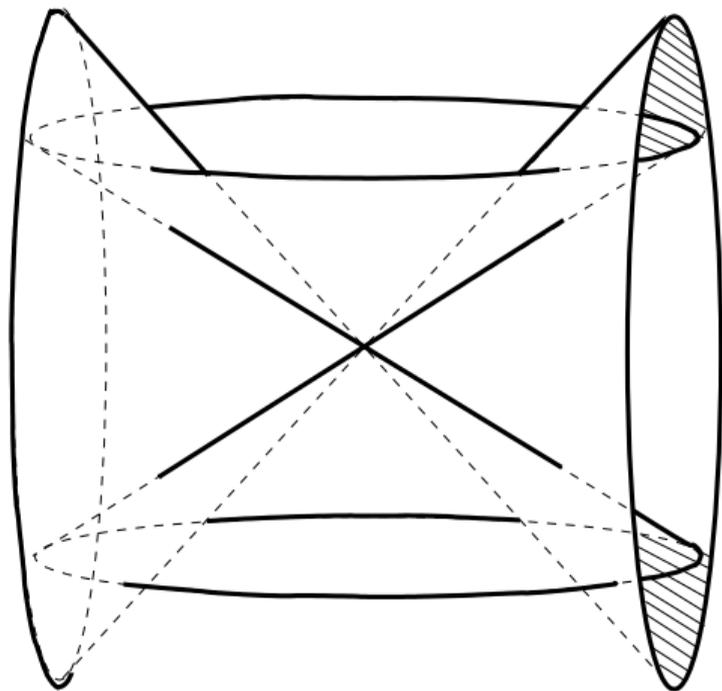
- Lastly, also tied to general/modified/...dispersion relations $\omega(\mathbf{p})$ or $P_x(q) = m^{\deg P}, P_x : T_x^* M \rightarrow \mathbb{R}$.⁷

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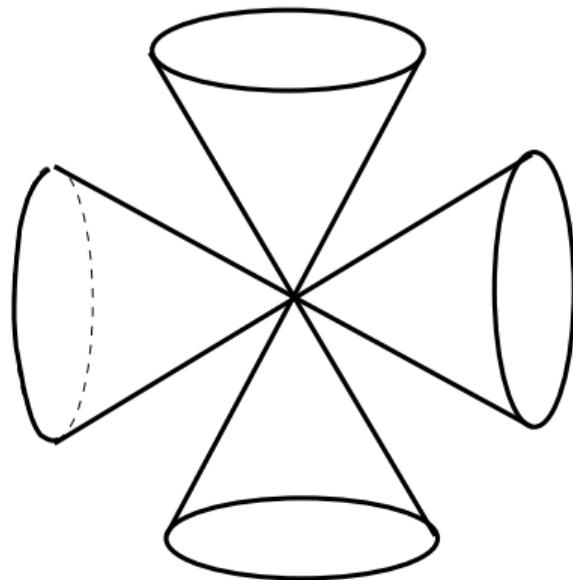
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⁷See Rätzel, Rivera, Schuller (arXiv:[1010.1369](https://arxiv.org/abs/1010.1369)). **Warning!** Technical & physical requirements on P omitted.

What It Looks Like



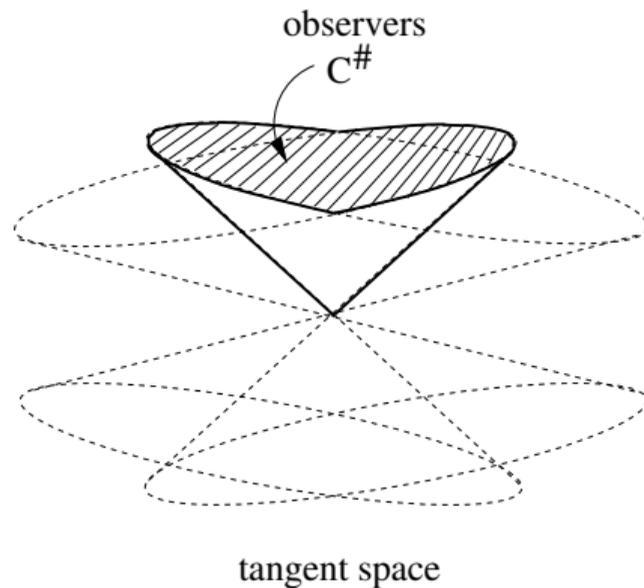
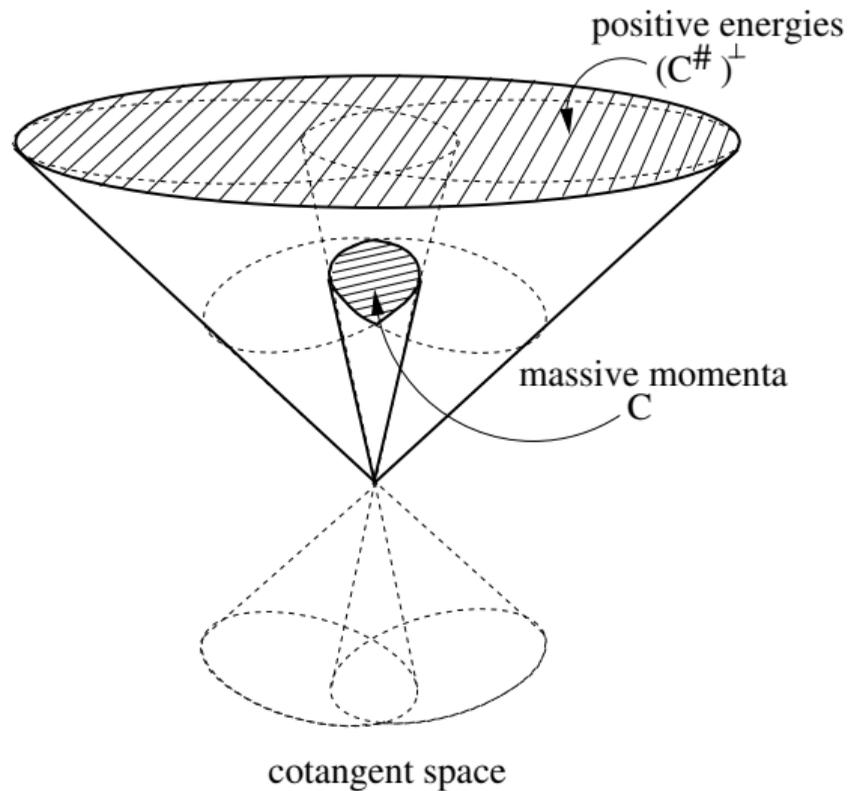
cotangent space



tangent space

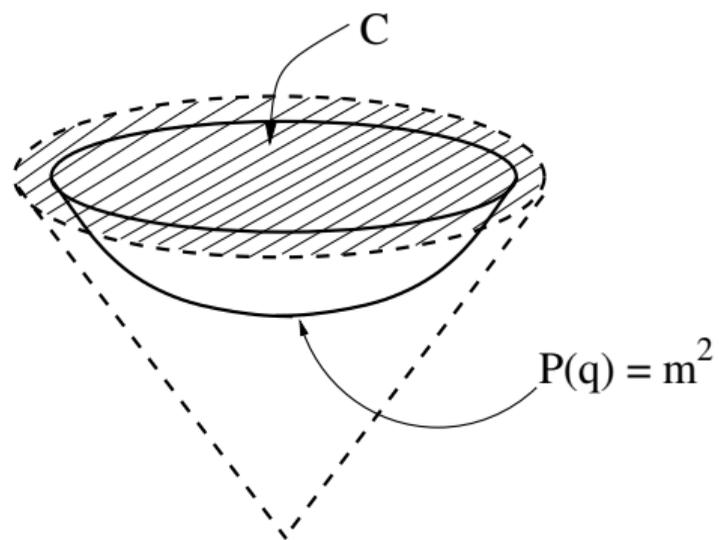
Images sourced from: Rätz, Rivera, Schuller (arXiv:[1010.1369](https://arxiv.org/abs/1010.1369))

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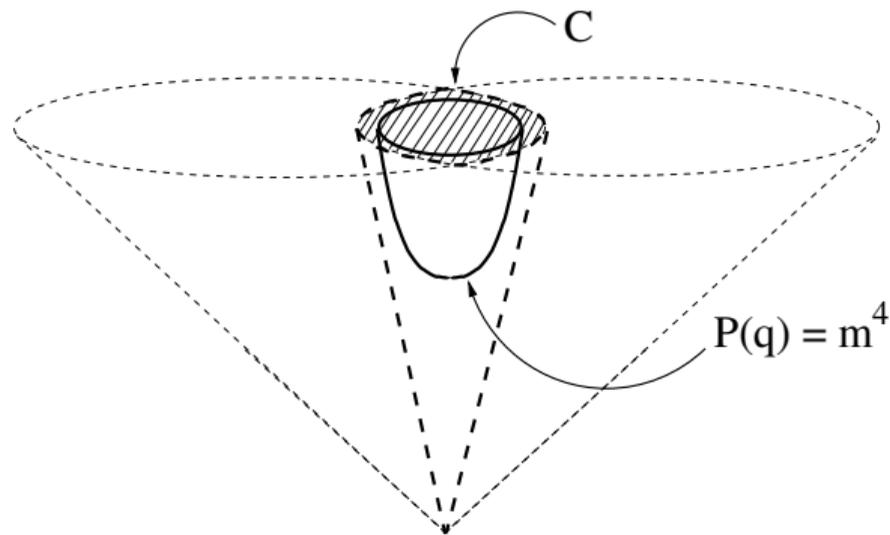


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cotangent space



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An Unruh–DeWitt Detector in a Uniaxial Crystal

The Detector Model

- Ingredients for uniaxial crystal:
 - Flat background[®] η ,
 - Crystal rest frame's worldline U ,
 - $Z^{abcd} = \eta^{a[c}\eta^{d]b} + U^{[a}X^{b]}U^{[c}X^{d]}$

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⁸Well, unless the constitutive tensor is assumed to *define* a more general (birefringent) space-time geometry.

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- $\eta(\dot{x}, \dot{x}) = \eta(U, U) = -1$, $\eta(M, \dot{x}) = \eta(U, X)$

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 - Dipole M^a
- $\eta(\dot{x}, \dot{x}) = \eta(U, U) = -1$, $\eta(M, \dot{x}) = \eta(U, X)$
- Fresnel equation⁹ determines wave front momenta: $\eta^{-1}(k, k)\zeta^{-1}(k, k) = 0$ with

$$\zeta_{ab} = \eta_{ab} + \frac{\xi^2}{1 + \xi^2} U_a U_b - \frac{1}{1 + \xi^2} X_a X_b$$

⁸Well, unless the constitutive tensor is assumed to *define* a more general (birefringent) space-time geometry.

⁹See, for example, Hehl, Obukhov – *Foundations of Classical Electrodynamics*

Early Results: Inertial Motion

- Building on results of Fewster, Pfeifer, Siemssen (arXiv:1709.01760, arXiv:1602.00946)
- Currently, only (partially) done for inertial \dot{x}
- Calculate $\mathcal{M}^a \mathcal{M}^c \dot{x}^b \dot{x}^d \langle 0 | F_{cd}(x(\tau')) F_{ab}(x(\tau)) | 0 \rangle \Big|_{\xi=0}$
- Using clever variable choices¹⁰

$$U^a = (1, \vec{0})^a, \hat{X}^a = (0, \hat{x})^a,$$

$$\dot{x}^a = \gamma(1, v)^a,$$

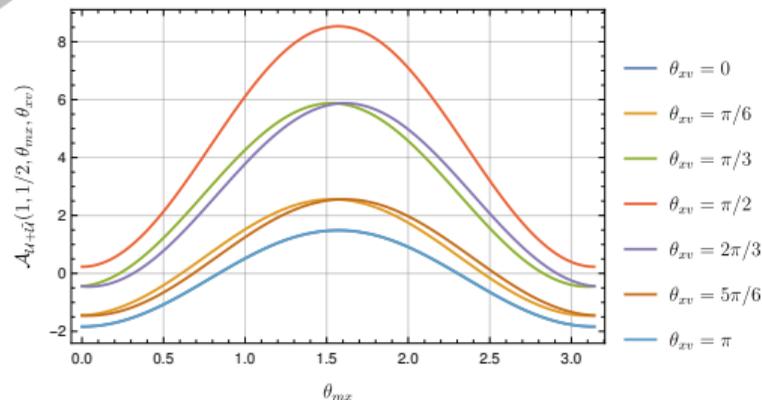
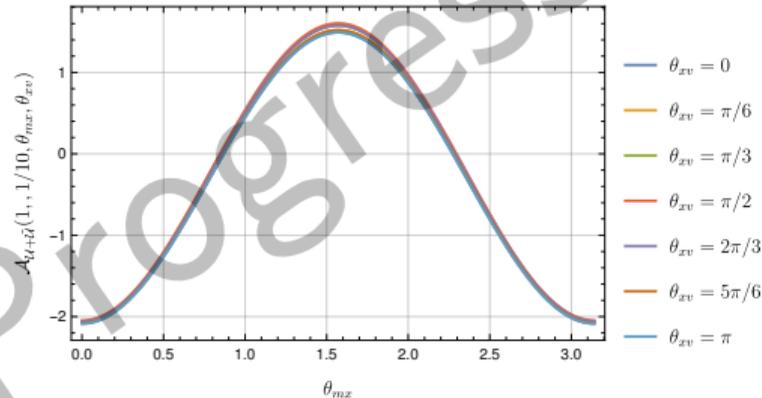
$$\eta(U, \dot{x}) = -\gamma$$

$$\mathcal{M}^a = |\mathcal{M}| (\gamma v \cos \theta_{mv} \hat{m} + (\gamma - 1) \cos \theta_{mv} \hat{v})^a,$$

$$\cos \theta_{mv} = \hat{m} \cdot \hat{v},$$

$$\eta(\hat{X}, \dot{x}) = v\gamma \cos \theta_{vx},$$

isolate factors $A_{U+\hat{U}}(\xi, v, \theta_{mx}, \theta_{mv})$



¹⁰Thank you, Finn!

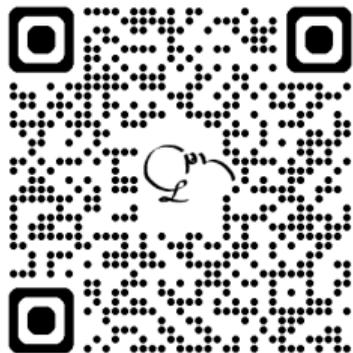
Summary

Summary:

- Analogues fit naturally into geometries beyond Lorentzian metrics
- One can find covariant dispersion relations
- Many application to modified theories of gravity

Outlook

- Interial UdW detector response in a uniaxial crystal
- Hopefully also accelerated UdW detector response...



From Black Holes to the Cosmos

Matt Visser's Journey through Space and Time



SISSA Miramare Campus
Trieste, Italy
24-28 August 2026

Speakers include:

- Carl Bender
- Ivan Booth
- Erik Curiel
- Fay Dowker
- Ted Jacobson
- Eleni Alexandra-Kontou
- Francisco Lobo
- Robert Mann
- Eric Poisson
- Ralf Schützhold
- Thomas Sotiriou
- Bill Unruh
- Cliff Will
- David Wiltshire



Information and Registration under:
<https://indico.sissa.it/event/175/>



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Electromagnetism: Microscopic vs. Macroscopic

Microscopic Maxwell equations

$$\begin{aligned}\nabla_{[a}F_{bc]} &= 0, \\ \nabla_a F^{ab} &= J^b\end{aligned}$$

non-trivial consti-
tutive relations

Macroscopic Maxwell equations

$$\begin{aligned}\nabla_{[a}F_{bc]} &= 0, \\ \nabla_a (\underbrace{Z^{abcd} F_{cd}}_{G^{ab}}) &= J^b,\end{aligned}$$

$$\mathbf{D} = \epsilon \mathbf{E} + \zeta \mathbf{B},$$

$$\mathbf{H} = \zeta^\dagger \mathbf{E} + \mu^{-1} \mathbf{B}.$$

What It Looks Like—Addendum ‘Observers’

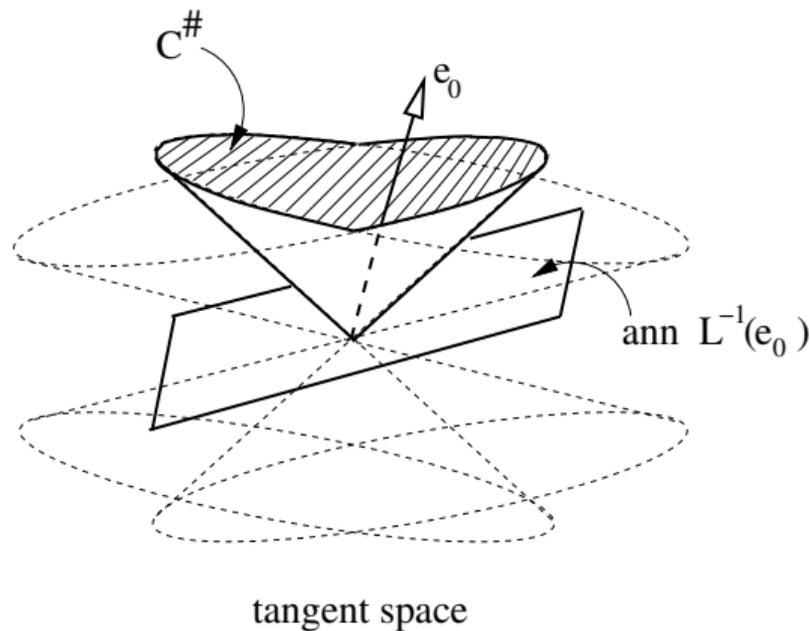
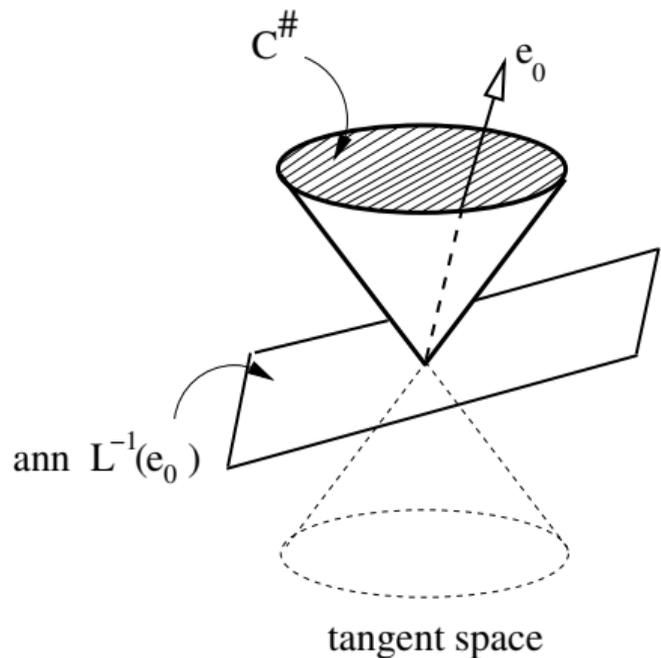


Image sourced from: Rätzel, Rivera, Schuller (arXiv:[1010.1369](https://arxiv.org/abs/1010.1369))

What It Looks Like—Addendum ‘Stability’

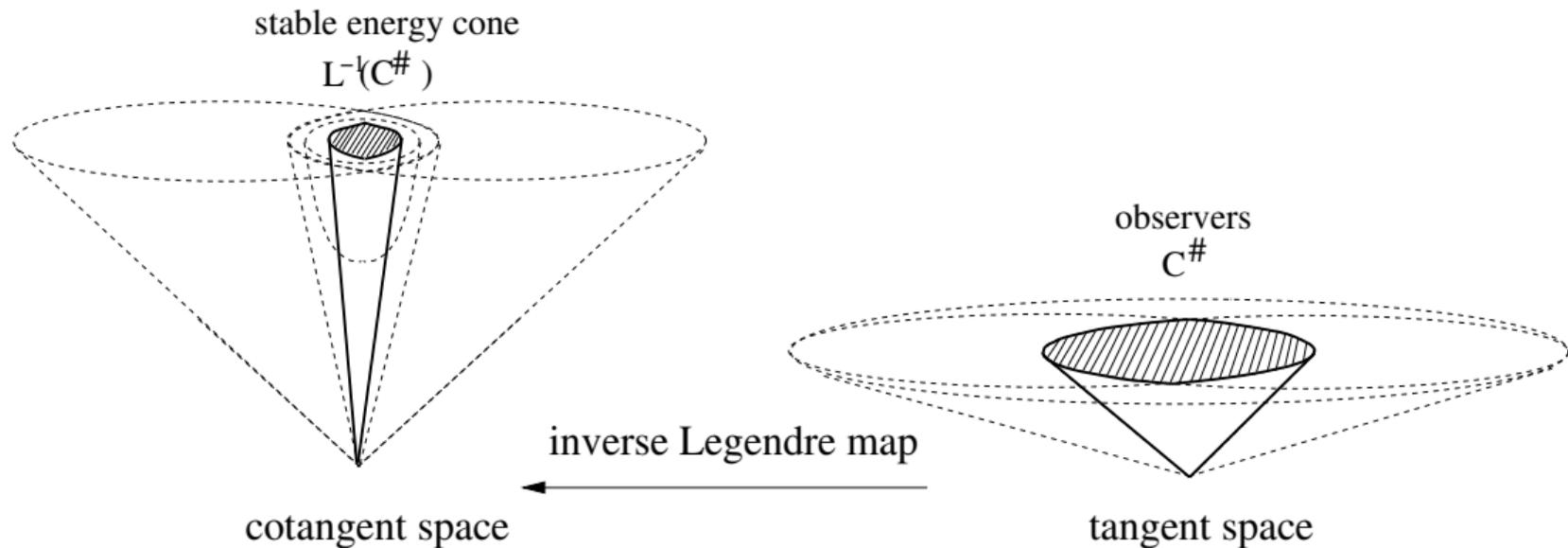


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