



BSM with LHCb: Theory overview of the implications of rare decays

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Outline

- Introduction and motivations
- Flavour observables
 - $B_{s,d} \rightarrow \mu^+ \mu^-$
 - $B \rightarrow K^* \mu^+ \mu^-$
- Implications for New Physics
 - Model independent constraints
 - Applications for Supersymmetry
- SuperIso
- Conclusion

Why rare decays?

- sensitivity to new physics effects
- complementary to other searches
- probe sectors inaccessible to direct searches
- test quantum structure of the SM at loop level
- constrain parameter spaces of new physics scenarios
- valuable data already available
- promising experimental situation
- consistency checks with direct observations

Flavour physics and rare decays in particular are excellent tools to probe BSM physics!

A multi-scale problem

- new physics: $1/\Lambda_{\text{NP}}$
- electroweak interactions: $1/M_W$
- hadronic effects: $1/m_b$
- QCD interactions: $1/\Lambda_{\text{QCD}}$

⇒ Effective field theory approach:

separation between low and high energies using Operator Product Expansion

- short distance: Wilson coefficients, computed perturbatively
- long distance: local operators

$$\mathcal{H}_{\text{eff}} = -\frac{4G_F}{\sqrt{2}} V_{tb} V_{ts}^* \left(\sum_{i=1 \dots 10, S, P} (C_i(\mu) \mathcal{O}_i(\mu) + C'_i(\mu) \mathcal{O}'_i(\mu)) \right)$$

New physics:

- Corrections to the Wilson coefficients: $C_i \rightarrow C_i + \delta C_i^{\text{NP}}$
- Additional operators: $\sum_j C_j^{\text{NP}} \mathcal{O}_j^{\text{NP}}$

Why flavour physics is complicated?

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Operators

$$\mathcal{O}_7 = \frac{e}{g^2} m_b (\bar{s} \sigma_{\mu\nu} P_R b) F^{\mu\nu}$$

$$\mathcal{O}_8 = \frac{1}{g} m_b (\bar{s} \sigma_{\mu\nu} T^a P_R b) G^{\mu\nu a}$$

$$\mathcal{O}_9 = \frac{e^2}{g^2} (\bar{s} \gamma_\mu P_L b) (\bar{\mu} \gamma^\mu \mu)$$

$$\mathcal{O}_{10} = \frac{e^2}{g^2} (\bar{s} \gamma_\mu P_L b) (\bar{\mu} \gamma^\mu \gamma_5 \mu)$$

$$\mathcal{O}_S = \frac{e^2}{16\pi^2} m_b (\bar{s} P_R b) (\bar{\mu} \mu)$$

$$\mathcal{O}_P = \frac{e^2}{16\pi^2} m_b (\bar{s} P_R b) (\bar{\mu} \gamma_5 \mu)$$

$$\mathcal{O}'_7 = \frac{e}{g^2} m_b (\bar{s} \sigma_{\mu\nu} P_L b) F^{\mu\nu}$$

$$\mathcal{O}'_8 = \frac{1}{g} m_b (\bar{s} \sigma_{\mu\nu} T^a P_L b) G^{\mu\nu a}$$

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Primed operators: opposite chirality to the unprimed ones,
vanish or highly suppressed in the SM

Wilson coefficients

Two main steps:

- Calculating $C_i^{\text{eff}}(\mu)$ at scale $\mu \sim M_W$ by requiring matching between the effective and full theories

$$C_i^{\text{eff}}(\mu) = C_i^{(0)\text{eff}}(\mu) + \frac{\alpha_s(\mu)}{4\pi} C_i^{(1)\text{eff}}(\mu) + \dots$$

- Evolving the $C_i^{\text{eff}}(\mu)$ to scale $\mu \sim m_b$ using the RGE:

$$\mu \frac{d}{d\mu} C_i^{\text{eff}}(\mu) = C_j^{\text{eff}}(\mu) \gamma_{ji}^{\text{eff}}(\mu)$$

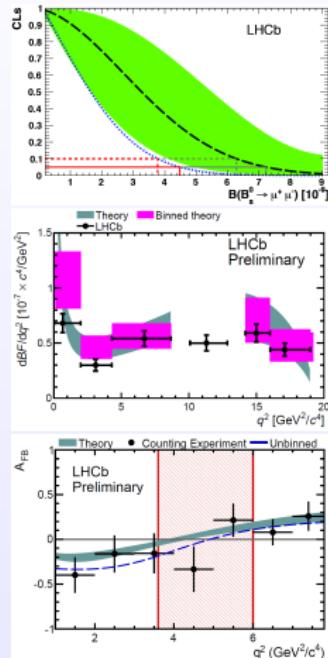
driven by the anomalous dimension matrix $\hat{\gamma}^{\text{eff}}(\mu)$:

$$\hat{\gamma}^{\text{eff}}(\mu) = \frac{\alpha_s(\mu)}{4\pi} \hat{\gamma}^{(0)\text{eff}} + \frac{\alpha_s^2(\mu)}{(4\pi)^2} \hat{\gamma}^{(1)\text{eff}} + \dots$$

Observables

LHCb recent results:

- $B_{s,d} \rightarrow \mu^+ \mu^-$
- $B \rightarrow K^* \mu^+ \mu^-$
 - $\text{BR}(B \rightarrow K^* \mu^+ \mu^-)$
 - $A_{FB}(B \rightarrow K^* \mu^+ \mu^-)$
 - $F_L(B \rightarrow K^* \mu^+ \mu^-)$
 - $S_3(B \rightarrow K^* \mu^+ \mu^-)$
 - $A_{Im}(B \rightarrow K^* \mu^+ \mu^-)$
 - $A_{FB0}(B \rightarrow K^* \mu^+ \mu^-)$



Other observables:

- $B \rightarrow X_{s,d} \gamma$
- $B \rightarrow X_s \mu^+ \mu^-$
- $B \rightarrow \tau \nu$
- ...

$$\text{BR}(B_s \rightarrow \mu^+ \mu^-)$$

$$\mathcal{H}_{\text{eff}} = -\frac{4G_F}{\sqrt{2}} V_{tb} V_{ts}^* \left[\sum_{i=1 \dots 10, S, P} (C_i(\mu) \mathcal{O}_i(\mu) + C'_i(\mu) \mathcal{O}'_i(\mu)) \right]$$

Relevant operators:

$$\mathcal{O}_{10} = \frac{e^2}{(4\pi)^2} (\bar{s}\gamma^\mu b_L)(\bar{\ell}\gamma_\mu\gamma_5\ell)$$

$$\mathcal{O}_S = \frac{e^2}{16\pi^2} (\bar{s}_L^\alpha b_R^\alpha)(\bar{\ell}\ell)$$

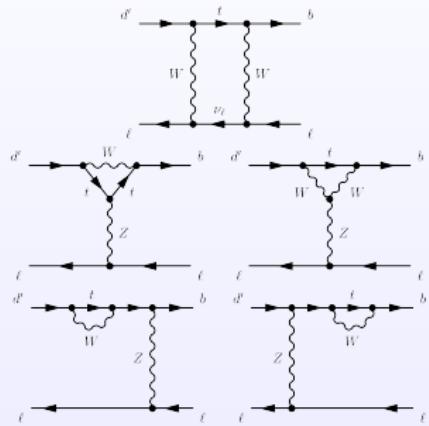
$$\mathcal{O}_P = \frac{e^2}{16\pi^2} (\bar{s}_L^\alpha b_R^\alpha)(\bar{\ell}\gamma_5\ell)$$

$$\text{BR}(B_s \rightarrow \mu^+ \mu^-) = \frac{G_F^2 \alpha^2}{64\pi^3} f_{B_s}^2 \tau_{B_s} m_{B_s}^3 |V_{tb} V_{ts}^*|^2 \sqrt{1 - \frac{4m_\mu^2}{m_{B_s}^2}}$$

$$\times \left\{ \left(1 - \frac{4m_\mu^2}{m_{B_s}^2} \right) |C_S - C'_S|^2 + \left| (C_P - C'_P) + 2(C_{10} - C'_{10}) \frac{m_\mu}{m_{B_s}} \right|^2 \right\}$$

Very sensitive to new physics, especially for large $\tan \beta$:

SUSY contributions can lead to an $O(100)$ enhancement over the SM!



Experimental results:

LHCb: $\text{BR}(B_s \rightarrow \mu^+ \mu^-) < 4.5 \times 10^{-9}$ at 95% C.L. [arXiv:1203.4493](https://arxiv.org/abs/1203.4493)CMS: $\text{BR}(B_s \rightarrow \mu^+ \mu^-) < 7.7 \times 10^{-9}$ at 95% C.L. [CMS BPH11020](https://cds.cern.ch/record/11020)

→ Approaching dangerously the SM value!

→ Crucial to have a clear estimation of the SM prediction!

Main source of uncertainty: f_{B_s}

- ETMC-11: 232 ± 10 MeV
- HPQCD-12:
 - HPQCD NR-09: 231 ± 15 MeV
 - HPQCD HISQ-11: 225 ± 4 MeV
- Fermilab-MILC-11: 242 ± 9.5 MeV

Our choice: 234 ± 10 MeV

$$\text{BR}(B_s \rightarrow \mu^+ \mu^-)$$

Most up-to-date input parameters (PDG 2011):

V_{ts}	V_{tb}	m_{B_s}	τ_{B_s}
-0.0403	0.999152	5.3663 GeV	1.472 ps

SM prediction: $\text{BR}(B_s \rightarrow \mu^+ \mu^-) = (3.53 \pm 0.38) \times 10^{-9}$

Most important sources of uncertainties:

8% from f_{B_s}

2% from B_s lifetime

2% from EW corrections

5% from V_{ts}

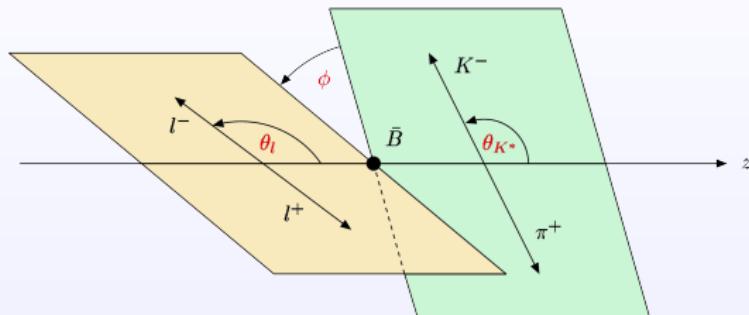
2% from scales

1.3% from top mass

Overall TH uncertainty: $\sim 10\%$.

Using $f_{B_s} = 225$ MeV and $\tau_{B_s} = 1.425$ ps, one gets: $\text{BR}(B_s \rightarrow \mu^+ \mu^-) = 3.20 \times 10^{-9}$

Angular distributions



The full angular distribution of the decay $\bar{B}^0 \rightarrow \bar{K}^{*0} \ell^+ \ell^-$ with $\bar{K}^{*0} \rightarrow K^- \pi^+$ on the mass shell is completely described by four independent kinematic variables:

- q^2 : dilepton invariant mass squared
- θ_ℓ : angle between ℓ^- and the \bar{B} in the dilepton frame
- θ_{K^*} : angle between K^- and \bar{B} in the $K^- \pi^+$ frame
- ϕ : angle between the normals of the $K^- \pi^+$ and the dilepton planes

$B \rightarrow K^* \mu^+ \mu^-$ – Differential decay distribution

Differential decay distribution:

$$\frac{d^4\Gamma}{dq^2 d \cos \theta_\ell d \cos \theta_{K^*} d\phi} = \frac{9}{32\pi} \textcolor{blue}{J}(q^2, \theta_\ell, \theta_{K^*}, \phi)$$

Kinematics: $4m_\ell^2 \leq q^2 \leq (M_B - m_{K^*})^2$, $-1 \leq \cos \theta_\ell \leq 1$, $-1 \leq \cos \theta_{K^*} \leq 1$, $0 \leq \phi \leq 2\pi$

$\textcolor{blue}{J}(q^2, \theta_\ell, \theta_{K^*}, \phi)$ are written in function of the angular coefficients $\textcolor{blue}{J}_{1-9}^{s,c}$

$\textcolor{blue}{J}_{1-9}$: functions of the spin amplitudes A_0 , $A_{||}$, A_{\perp} , A_t , and A_S

Spin amplitudes: functions of Wilson coefficients and form factors

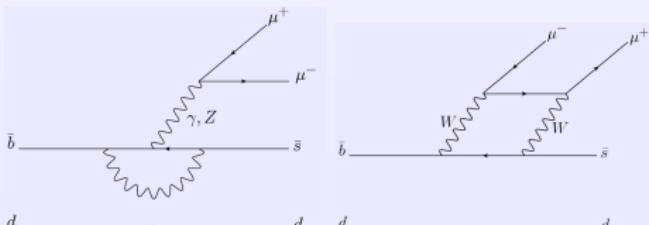
Main operators:

$$\mathcal{O}_9 = \frac{e^2}{(4\pi)^2} (\bar{s}\gamma^\mu b_L)(\bar{\ell}\gamma_\mu \ell)$$

$$\mathcal{O}_{10} = \frac{e^2}{(4\pi)^2} (\bar{s}\gamma^\mu b_L)(\bar{\ell}\gamma_\mu \gamma_5 \ell)$$

$$\mathcal{O}_S = \frac{e^2}{16\pi^2} (\bar{s}_L^\alpha b_R^\alpha)(\bar{\ell}\ell)$$

$$\mathcal{O}_P = \frac{e^2}{16\pi^2} (\bar{s}_L^\alpha b_R^\alpha)(\bar{\ell}\gamma_5 \ell)$$



Dilepton invariant mass spectrum

$$\frac{d\Gamma}{dq^2} = \frac{3}{4} \left(J_1 - \frac{J_2}{3} \right)$$

Forward backward asymmetry

Difference between the differential branching fractions in the forward and backward directions:

$$A_{FB}(q^2) \equiv \left[\int_{-1}^0 - \int_0^1 \right] d \cos \theta_I \frac{d^2 \Gamma}{dq^2 d \cos \theta_I} \Bigg/ \frac{d\Gamma}{dq^2} = \frac{3}{8} J_6 \Bigg/ \frac{d\Gamma}{dq^2}$$

→ Reduced theoretical uncertainty

Forward backward asymmetry zero-crossing

→ Reduced form factor uncertainties

$$q_0^2 \simeq -2m_b m_B \frac{C_9^{\text{eff}}(q_0^2)}{C_7} + O(\alpha_s, \Lambda/m_b)$$

→ fix the sign of C_9/C_7

Polarization fractions:

$$F_L(q^2) = \frac{|A_0|^2}{|A_0|^2 + |A_{\parallel}|^2 + |A_{\perp}|^2}$$

$$F_T(q^2) = 1 - F_L(q^2) = \frac{|A_{\perp}|^2 + |A_{\parallel}|^2}{|A_0|^2 + |A_{\parallel}|^2 + |A_{\perp}|^2}$$

K^* polarization parameter:

$$\alpha_{K^*}(q^2) = \frac{2F_L}{F_T} - 1 = \frac{2|A_0|^2}{|A_{\parallel}|^2 + |A_{\perp}|^2} - 1$$

Transverse asymmetries:

$$A_T^{(1)}(q^2) = \frac{-2\Re(A_{\parallel} A_{\perp}^*)}{|A_{\perp}|^2 + |A_{\parallel}|^2}$$

$$A_T^{(2)}(q^2) = \frac{|A_{\perp}|^2 - |A_{\parallel}|^2}{|A_{\perp}|^2 + |A_{\parallel}|^2}$$

$$A_T^{(3)}(q^2) = \frac{|A_{0L} A_{\parallel L}^* + A_{0R}^* A_{\parallel R}|}{\sqrt{|A_0|^2 |A_{\perp}|^2}}$$

$$A_T^{(4)}(q^2) = \frac{|A_{0L} A_{\perp L}^* - A_{0R}^* A_{\perp R}|}{|A_{0L} A_{\parallel L}^* + A_{0R}^* A_{\parallel R}|}$$

$$A_{Im}(q^2) = -2 \operatorname{Im} \left(\frac{A_{\parallel} A_{\perp}^*}{|A_{\perp}|^2 + |A_{\parallel}|^2} \right)$$

$$S_3(q^2) = \frac{1}{2} (1 - F_L(q^2)) A_T^{(2)}(q^2)$$

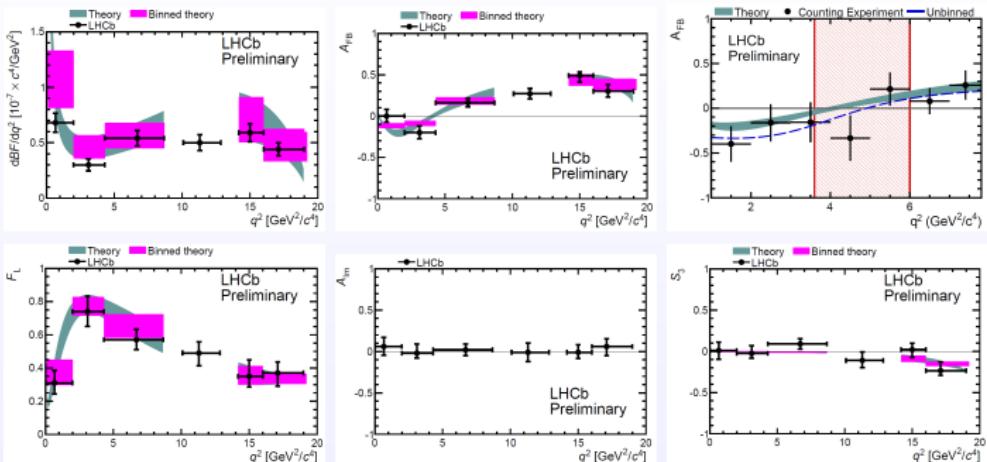
$B \rightarrow K^* \mu^+ \mu^-$ – SM predictions

Observable	SM value	(FF)	(SL)	(QM)	(CKM)	(Scale)
$10^7 \times BR(B \rightarrow K^* \mu^+ \mu^-)_{[1,6]}$	2.32	± 1.34	± 0.04	$+0.04$ -0.03	$+0.08$ -0.13	$+0.09$ -0.05
$\langle A_{FB}(B \rightarrow K^* \mu^+ \mu^-) \rangle_{[1,6]}$	-0.06	± 0.04	± 0.02	± 0.01	—	—
$\langle F_L(B \rightarrow K^* \mu^+ \mu^-) \rangle_{[1,6]}$	0.71	± 0.13	± 0.01	± 0.01	—	—
$q_0^2(B \rightarrow K^* \mu^+ \mu^-)/\text{GeV}^2$	4.26	± 0.30	± 0.15	$+0.14$ -0.04	—	$+0.02$ -0.04

Main uncertainties from:

- form factors
- $1/m_b$ subleading corrections
- parametric uncertainties (m_b , m_c , m_t)
- CKM matrix elements
- scales

$B \rightarrow K^* \mu^+ \mu^-$ – Experimental results from LHCb



q^2 range (GeV $^2/c^4$)	$dB\bar{F}/dq^2$ ($\times 10^{-7}$ GeV $^{-2}c^4$)	A_{FB}	F_L	A_{Im}	$2S_3$
$0.05 < q^2 < 2.00$	$0.68 \pm 0.07 \pm 0.05$	$0.00^{+0.08+0.01}_{-0.07-0.01}$	$0.31^{+0.07+0.03}_{-0.06-0.03}$	$0.06^{+0.11+0.00}_{-0.10-0.03}$	$0.02^{+0.20+0.00}_{-0.21-0.03}$
$2.00 < q^2 < 4.30$	$0.30 \pm 0.05 \pm 0.02$	$-0.20^{+0.08+0.01}_{-0.07-0.03}$	$0.74^{+0.09+0.02}_{-0.08-0.04}$	$-0.02^{+0.10+0.05}_{-0.06-0.01}$	$-0.05^{+0.18+0.05}_{-0.12-0.01}$
$4.30 < q^2 < 8.68$	$0.54 \pm 0.05 \pm 0.05$	$0.16^{+0.05+0.01}_{-0.05-0.01}$	$0.57^{+0.05+0.04}_{-0.05-0.03}$	$0.02^{+0.07+0.01}_{-0.07-0.01}$	$0.18^{+0.13+0.01}_{-0.13-0.01}$
$10.09 < q^2 < 12.89$	$0.50 \pm 0.06 \pm 0.04$	$0.27^{+0.06+0.02}_{-0.06-0.01}$	$0.49^{+0.06+0.03}_{-0.07-0.03}$	$-0.01^{+0.11+0.02}_{-0.11-0.03}$	$-0.22^{+0.20+0.02}_{-0.17-0.03}$
$14.18 < q^2 < 16.00$	$0.59 \pm 0.07 \pm 0.04$	$0.49^{+0.04+0.02}_{-0.06-0.05}$	$0.35^{+0.07+0.07}_{-0.06-0.02}$	$-0.01^{+0.08+0.04}_{-0.07-0.02}$	$0.04^{+0.15+0.04}_{-0.19-0.02}$
$16.00 < q^2 < 19.00$	$0.44 \pm 0.05 \pm 0.03$	$0.30^{+0.07+0.04}_{-0.07-0.01}$	$0.37^{+0.06+0.03}_{-0.07-0.04}$	$0.06^{+0.09+0.03}_{-0.10-0.05}$	$-0.47^{+0.21+0.03}_{-0.10-0.05}$
$1.00 < q^2 < 6.00$	$0.42 \pm 0.04 \pm 0.04$	$-0.18^{+0.06+0.01}_{-0.06-0.02}$	$0.66^{+0.06+0.04}_{-0.06-0.03}$	$0.07^{+0.07+0.02}_{-0.07-0.01}$	$0.10^{+0.15+0.02}_{-0.16-0.01}$

LHCb-CONF-2012-008

Implications

Assuming Minimal Flavour Violation (MFV)

What are the presently allowed ranges of the Wilson coefficients?

Operators of interest:

$$\mathcal{O}_7, \mathcal{O}_8, \mathcal{O}_9, \mathcal{O}_{10} \quad \text{and} \quad \mathcal{O}_{S-P} \propto (\bar{s}P_R b)(\bar{\mu}P_L \mu) \equiv \mathcal{O}'_0$$

NP manifests itself in the shifts of the individual coefficients with respect to the SM values:

$$C_i(\mu) = C_i^{\text{SM}}(\mu) + \delta C_i$$

- Scans over the values of $\delta C_7, \delta C_8, \delta C_9, \delta C_{10}, \delta C'_0$
- Calculation of flavour observables
- Comparison with experimental results
- Constraints on the Wilson coefficients C_i

see also: Hurth, Isidori, Kamenik, Mescia, Nucl.Phys. B808 (2009) 326
Descotes-Genon, Gosh, Matias, Ramon, JHEP 1106 (2011) 099
Altmannshofer, Paradisi, Straub, JHEP 1204 (2012) 008

Model independent constraints on New Physics

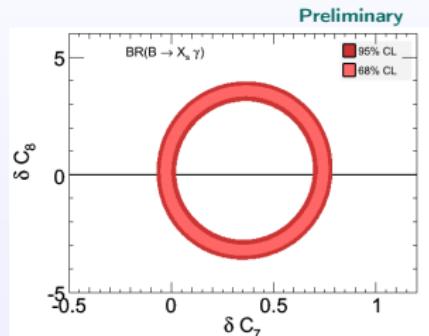
→ Global fits of the $\Delta F = 1$ observables obtained by minimization of

$$\chi^2 = \sum_i \frac{(O_i^{\text{exp}} - O_i^{\text{th}})^2}{(\sigma_i^{\text{exp}})^2 + (\sigma_i^{\text{th}})^2}$$

Observables:

- $\text{BR}(B \rightarrow X_s \gamma)$
- $\text{BR}(B \rightarrow X_d \gamma)$
- $\Delta_0(B \rightarrow K^* \gamma)$
- $\text{BR}^{\text{low}}(B \rightarrow X_s \mu^+ \mu^-)$
- $\text{BR}^{\text{high}}(B \rightarrow X_s \mu^+ \mu^-)$
- $\text{BR}(B_s \rightarrow \mu^+ \mu^-)$
- $\text{BR}^{\text{low}}(B \rightarrow K^* \mu^+ \mu^-)$
- $\text{BR}^{\text{high}}(B \rightarrow K^* \mu^+ \mu^-)$
- $A_{FB}^{\text{low}}(B \rightarrow K^* \mu^+ \mu^-)$
- $A_{FB}^{\text{high}}(B \rightarrow K^* \mu^+ \mu^-)$
- $q_0^2(A_{FB}(B \rightarrow K^* \mu^+ \mu^-))$
- $F_L^{\text{low}}(B \rightarrow K^* \mu^+ \mu^-)$

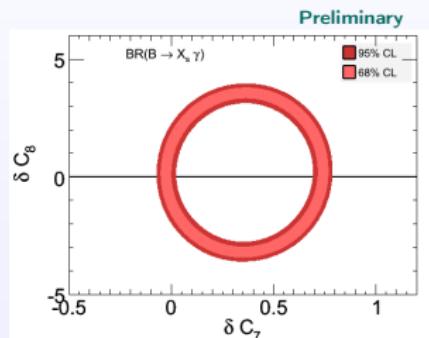
$B \rightarrow X_s \gamma$: sensitive to C_7 and C_8



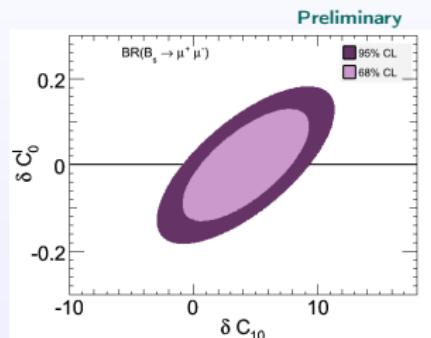
- No linear combination assumed for NP contributions to the electromagnetic and chromomagnetic operators
- Scalar operator strongly restricted by the $\text{BR}(B_s \rightarrow \mu^+ \mu^-)$ constraint

Model independent constraints on New Physics: individual observables

$B \rightarrow X_s \gamma$: sensitive to C_7 and C_8



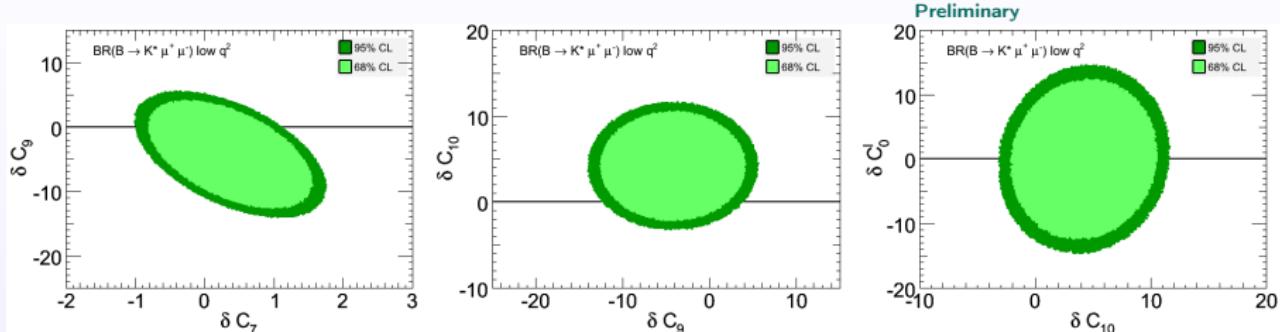
$B_s \rightarrow \mu^+ \mu^-$: sensitive to C_{10} and C'_0



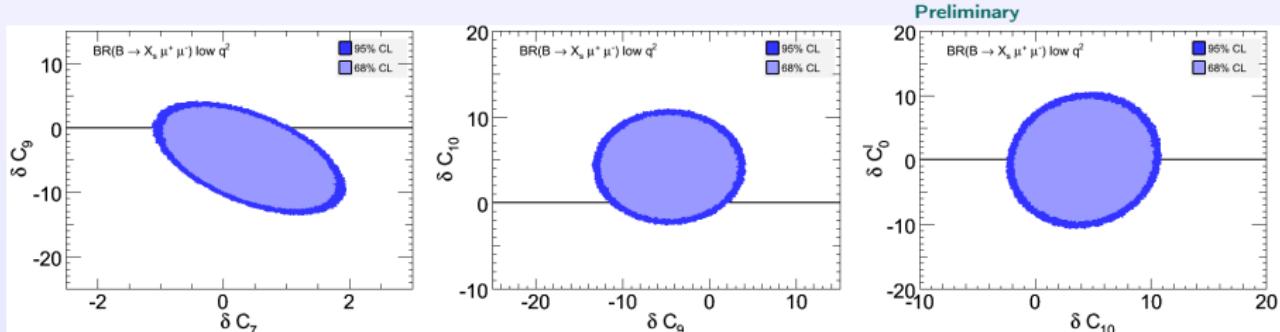
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Model independent constraints on New Physics: individual observables

$B \rightarrow K^* \mu^+ \mu^-$ exclusive mode

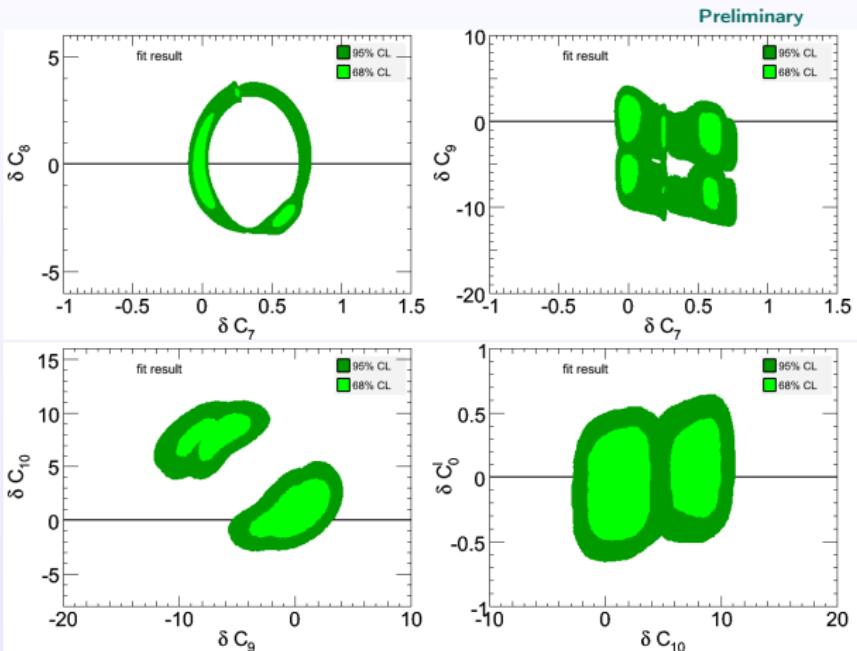


$B \rightarrow X_s \mu^+ \mu^-$ inclusive mode



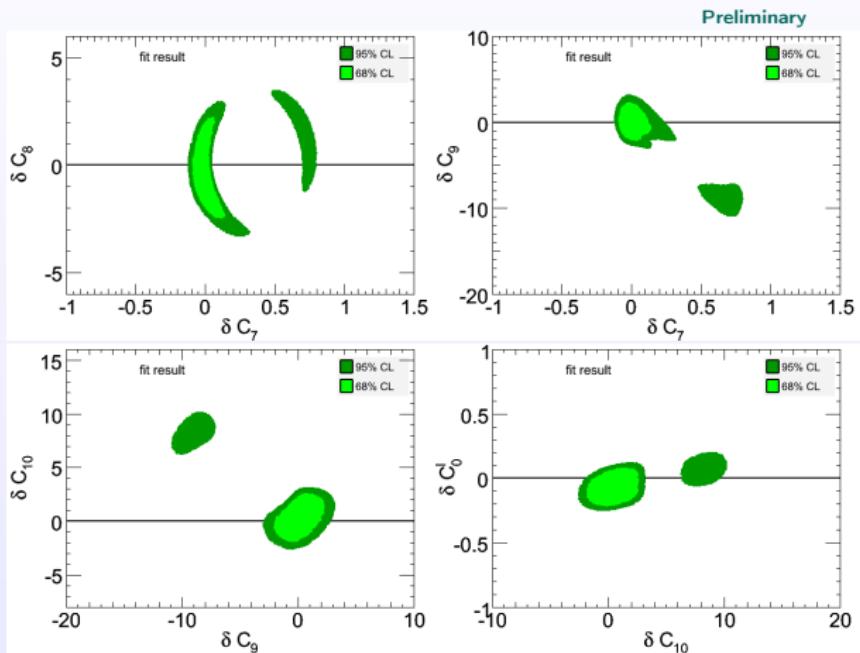
Model independent constraints on New Physics: fits

Before LHCb:



Model independent constraints on New Physics: fits

After LHCb:



Model independent constraints on New Physics: predictions

Use the allowed ranges for the Wilson coefficients to make predictions for the observables which are not yet measured

In particular:

- $\text{BR}(B_d \rightarrow \mu^+ \mu^-) < 0.32 \times 10^{-9}$

Current LHCb limit: $\text{BR}(B_d \rightarrow \mu^+ \mu^-) < 1.0 \times 10^{-9}$

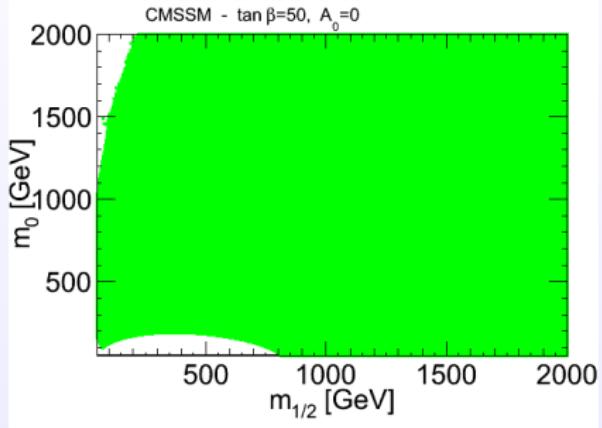
- $B \rightarrow K^* \mu^+ \mu^-$ transverse asymmetries:

- $A_T^{(2)} \in [-0.068, -0.02]$
- $A_T^{(3)} \in [0.35, 1.00]$
- $A_T^{(4)} \in [0.18, 1.30]$
- $A_T^{(5)} \in [0.15, 0.49]$

→ **Test of the MFV hypothesis!**

Constraints on Supersymmetry

Constraints on CMSSM

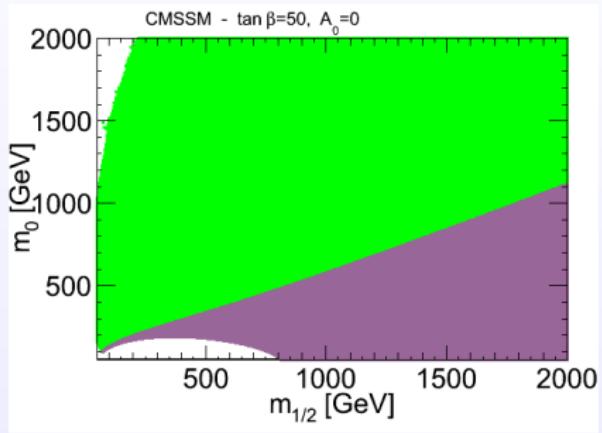


SuperIso v3.3

CMSSM parameter space

$$\tan \beta = 50, A_0 = 0$$

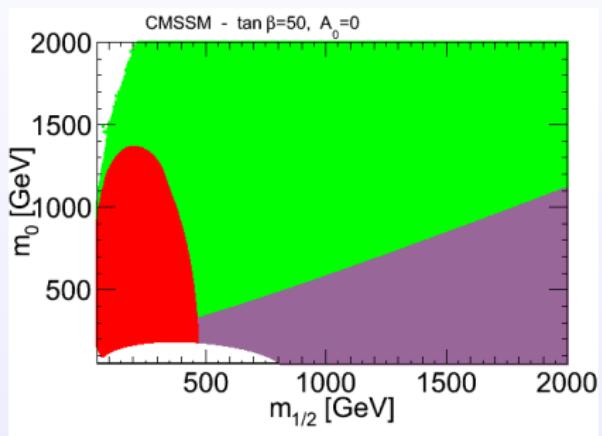
Constraints on CMSSM



SuperIso v3.3

Stable charged LSP incompatible with cosmology

Constraints on CMSSM



SuperIso v3.3

$$B \rightarrow X_s \gamma$$

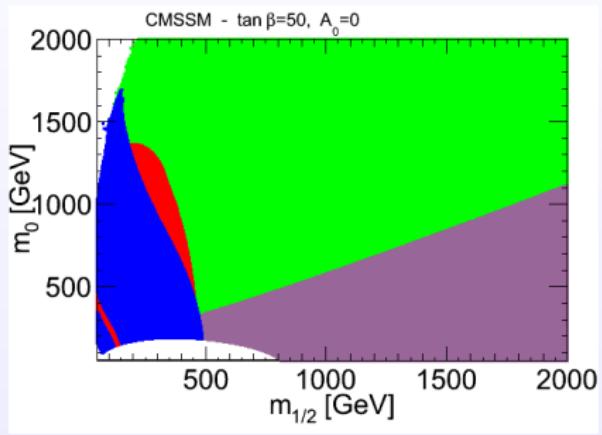
NNLO calculation

Misiak et al., PRL 98, 022002

Experimental value – HFAG 2011:

$$\text{BR}(\bar{B} \rightarrow X_s \gamma) = (3.55 \pm 0.25) \times 10^{-4}$$

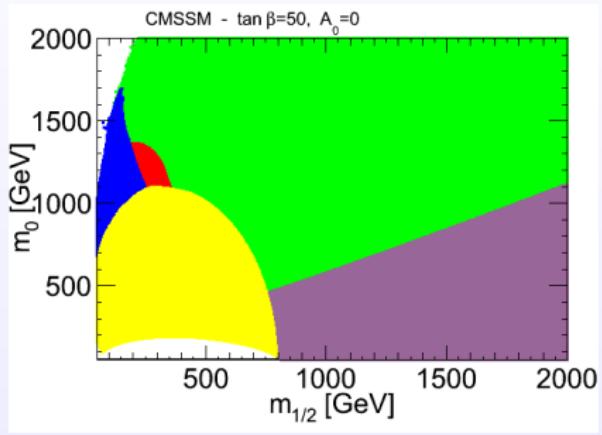
Constraints on CMSSM



$$\text{BR}(B \rightarrow \tau \nu) = \frac{G_F^2 |V_{ub}|^2}{8\pi} m_\tau^2 f_B^2 m_B \\ \times \left(1 - \frac{m_\tau^2}{m_B^2}\right)^2 \left|1 - \left(\frac{m_B^2}{m_{H^+}^2}\right) \frac{\tan^2 \beta}{1 + \epsilon_0 \tan \beta}\right|^2$$

HFAG 2011:
 $\text{BR}(B \rightarrow \tau \nu) = (1.64 \pm 0.34) \times 10^{-4}$

Constraints on CMSSM

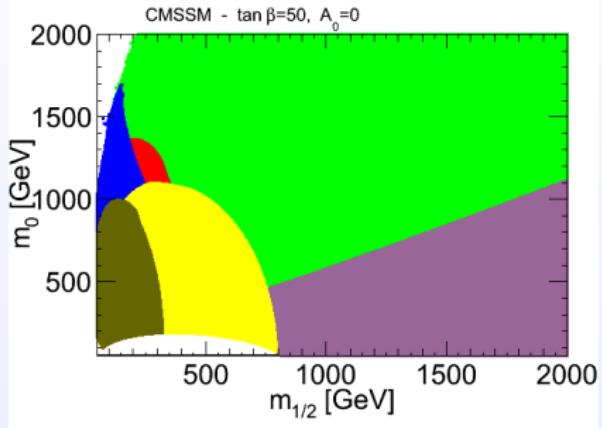


$$B_s \rightarrow \mu^+ \mu^-$$

EPS LHCb + CMS combination:

$$\text{BR}(B_s \rightarrow \mu^+ \mu^-) < 1.1 \times 10^{-8} \text{ at 95\% C.L.}$$

Constraints on CMSSM

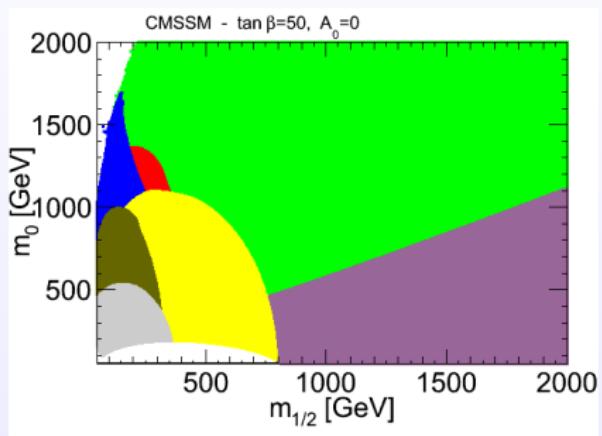


$$B \rightarrow K^* \mu^+ \mu^-$$

In the region $1 < q^2 < 6$ GeV 2 (LHCb):

$$\langle A_{FB}(B \rightarrow K^* \mu^+ \mu^-) \rangle = -0.18 \pm 0.06 \pm 0.02$$

Constraints on CMSSM



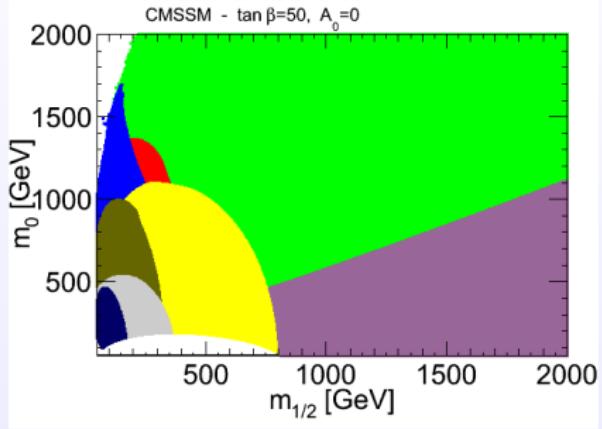
SuperIso v3.3

$$B_d \rightarrow \mu^+ \mu^-$$

EPS LHCb limit:

$$\text{BR}(B_d \rightarrow \mu^+ \mu^-) < 5.1 \times 10^{-9} \text{ at 95\% C.L.}$$

Constraints on CMSSM

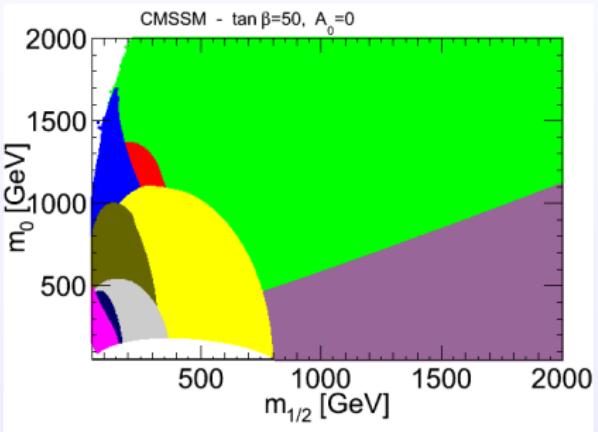


$$B \rightarrow K^* \mu^+ \mu^-$$

In the region $1 < q^2 < 6$ GeV 2 (LHCb):

$$\left\langle \frac{d\text{BR}}{dq^2}(B \rightarrow K^* \mu^+ \mu^-) \right\rangle = (0.42 \pm 0.04 \pm 0.04) \times 10^{-7}$$

Constraints on CMSSM

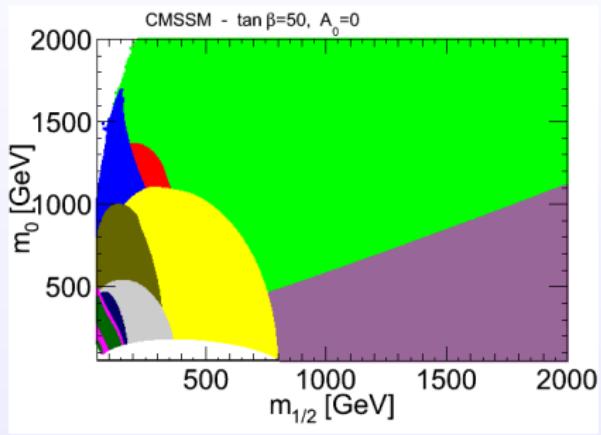


$K \rightarrow \mu\nu$

$$R_{\ell 23} = \left| \frac{V_{us}(K_{\ell 2})}{V_{us}(K_{\ell 3})} \times \frac{V_{ud}(0^+ \rightarrow 0^+)}{V_{ud}(\pi_{\ell 2})} \right| = \\ \left| 1 - \frac{m_{K^+}^2}{M_{H^+}^2} \left(1 - \frac{m_d}{m_s} \right) \frac{\tan^2 \beta}{1 + \epsilon_0 \tan \beta} \right|$$

Combined experimental and SM constraint:
 $R_{\ell 23}^{exp/SM} = 1.004 \pm 0.014$

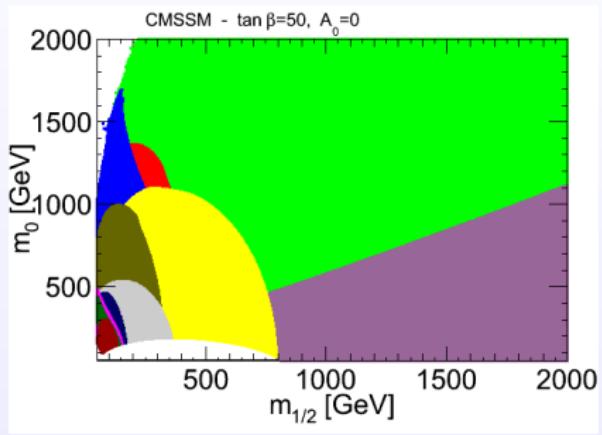
Constraints on CMSSM



$$\frac{d\Gamma(B \rightarrow D \ell \bar{\nu})}{dw} = \frac{G_F^2 |V_{cb}|^2 m_B^5}{192\pi^3} \rho_V(w) \times \\ \left[1 - \frac{m_\ell^2}{m_B^2} \left| 1 - \frac{t(w)}{(m_b - m_c)} \frac{m_b}{m_{H^+}^2} \frac{\tan^2 \beta}{1 + \epsilon_0 \tan \beta} \right|^2 \rho_S(w) \right]$$

PDG 2011:
$$\frac{\text{BR}(B^- \rightarrow D^0 \tau^- \bar{\nu})}{\text{BR}(B^- \rightarrow D^0 e^- \bar{\nu})} = 0.416 \pm 0.117 \pm 0.052$$

Constraints on CMSSM

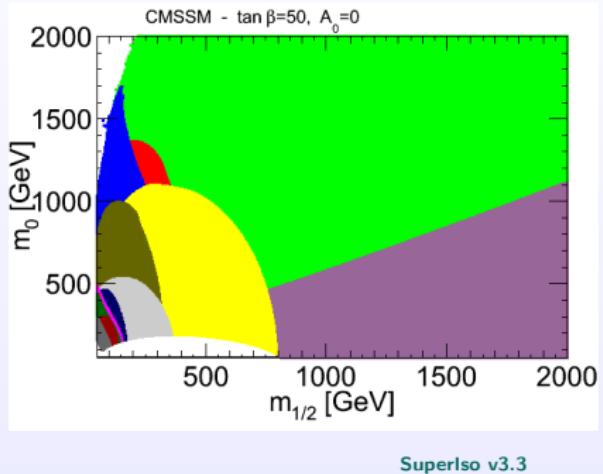


$B \rightarrow X_s \mu^+ \mu^-$, high q^2

Experimental value:

$$\text{BR}(B \rightarrow X_s \mu^+ \mu^-) = (4.18 \pm 1.35) \times 10^{-7}$$

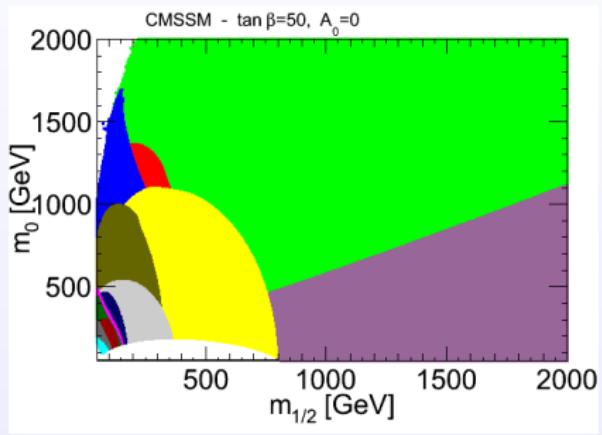
Constraints on CMSSM



$$\begin{aligned} \text{BR}(D_s \rightarrow \ell \nu) = & \\ \frac{G_F^2}{8\pi} |V_{cs}|^2 f_{D_s}^2 m_\ell^2 M_{D_s} \tau_{D_s} & \left(1 - \frac{m_\ell^2}{M_{D_s}^2}\right)^2 \\ \times \left[1 + \left(\frac{1}{m_c + m_s}\right) \left(\frac{M_{D_s}}{m_{H^+}}\right)^2\right. \\ \left. \times \left(m_c - \frac{m_s \tan^2 \beta}{1 + \epsilon_0 \tan \beta}\right)\right]^2 \end{aligned}$$

HFAG 2011:
 $\text{BR}(D_s \rightarrow \tau \nu) = (5.38 \pm 0.32) \times 10^{-2}$

Constraints on CMSSM

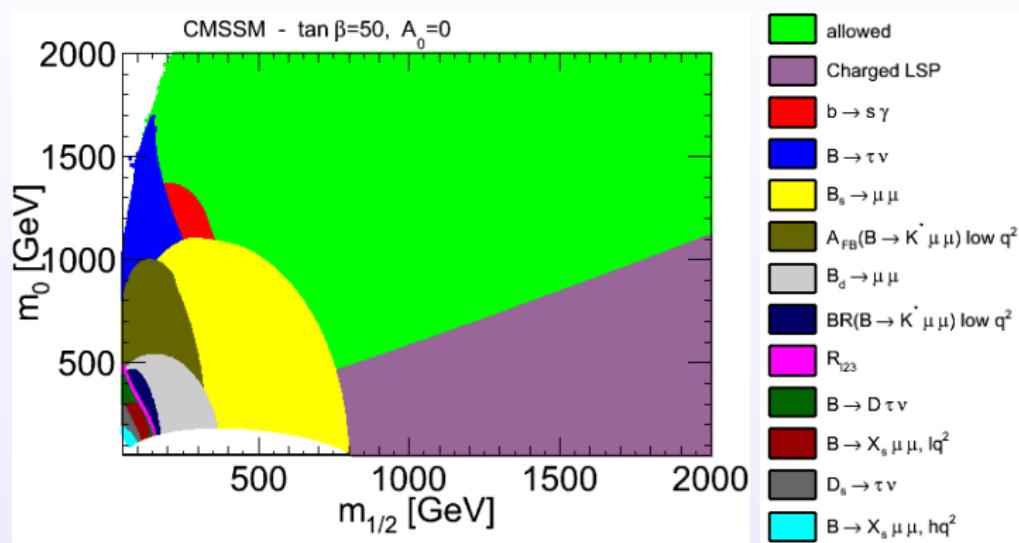


$B \rightarrow X_s \mu^+ \mu^-$, low q^2

Experimental value:

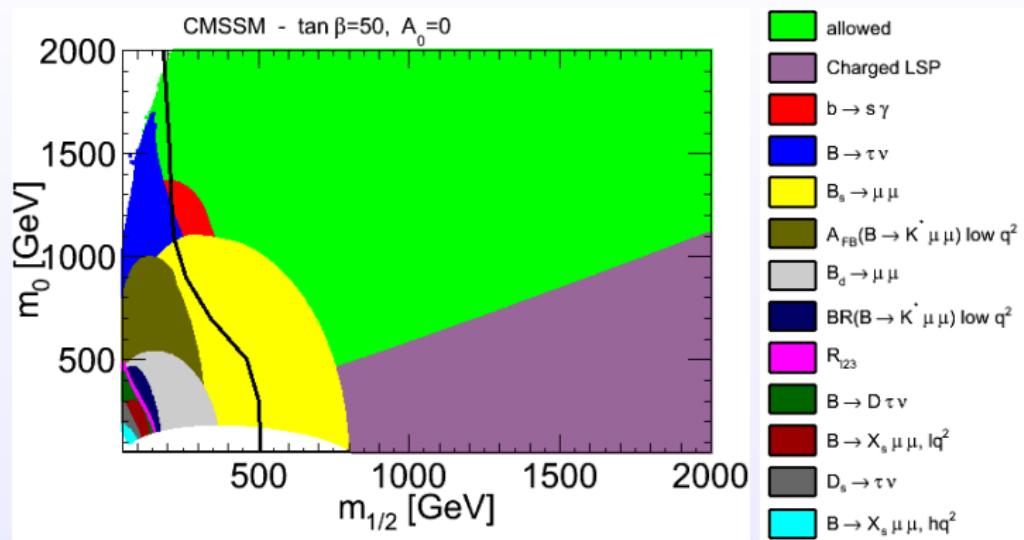
$$\text{BR}(B \rightarrow X_s \mu^+ \mu^-) = (1.60 \pm 0.68) \times 10^{-6}$$

Constraints on CMSSM



SuperIso v3.3

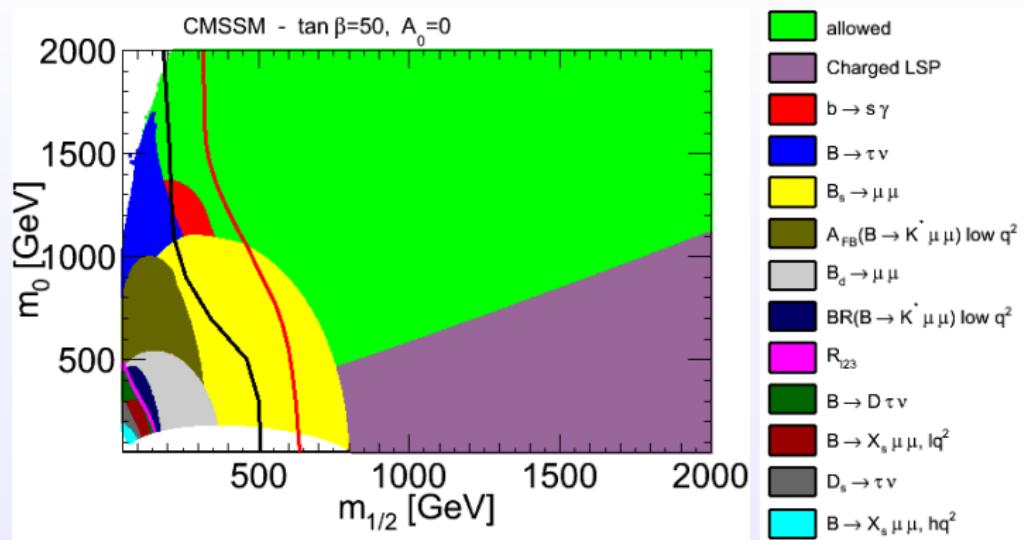
Constraints on CMSSM



Black line: CMS exclusion limit with 1.1 fb^{-1} data

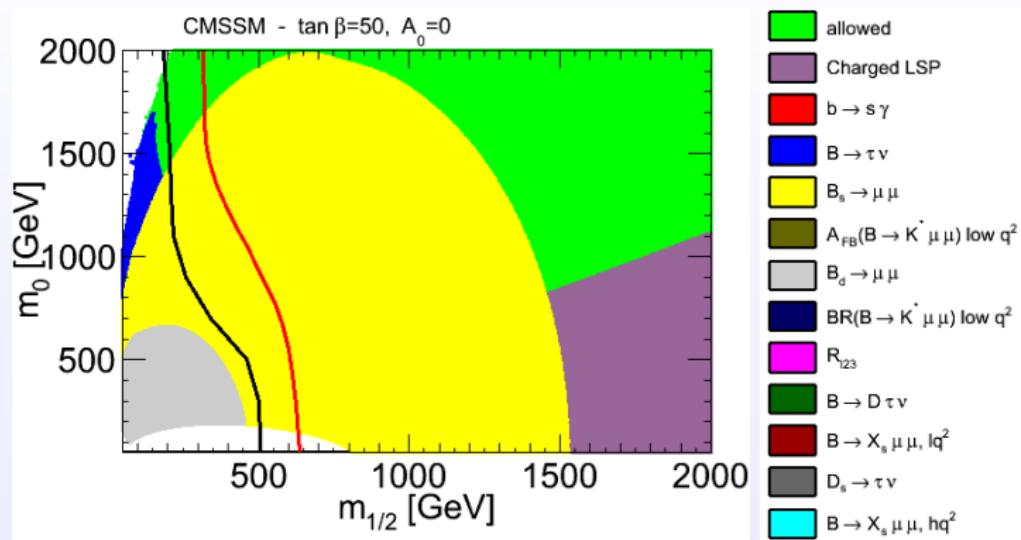
SuperIso v3.3

Constraints on CMSSM



Black line: CMS exclusion limit with 1.1 fb^{-1} data
Red line: CMS exclusion limit with 4.4 fb^{-1} data

Constraints on CMSSM



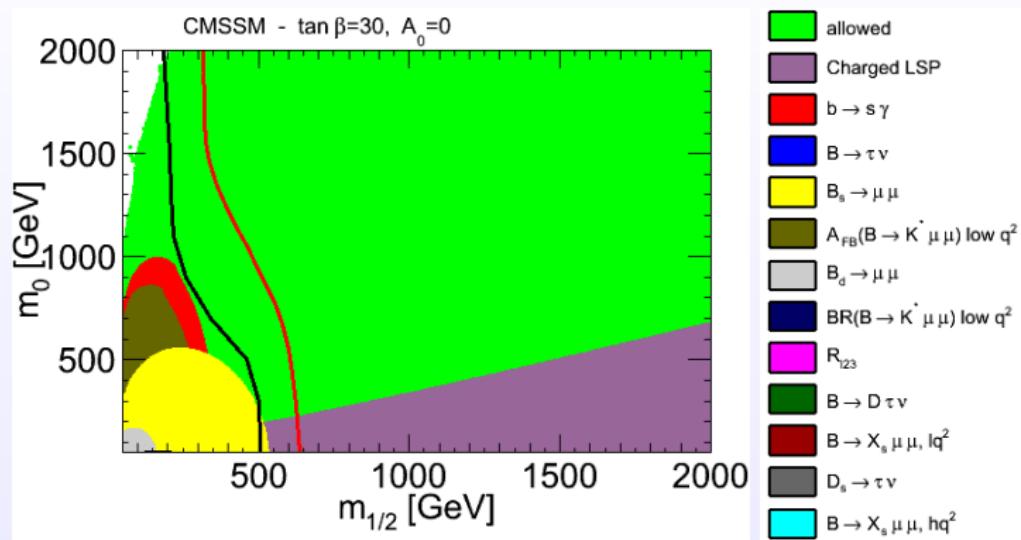
Black line: CMS exclusion limit with 1.1 fb^{-1} data

Red line: CMS exclusion limit with 4.4 fb^{-1} data

New LHCb limits for $BR(B_s \rightarrow \mu^+ \mu^-)$ and $BR(B_d \rightarrow \mu^+ \mu^-)$

SuperIso v3.3

Constraints on CMSSM



Black line: CMS exclusion limit with 1.1 fb^{-1} data

Red line: CMS exclusion limit with 4.4 fb^{-1} data

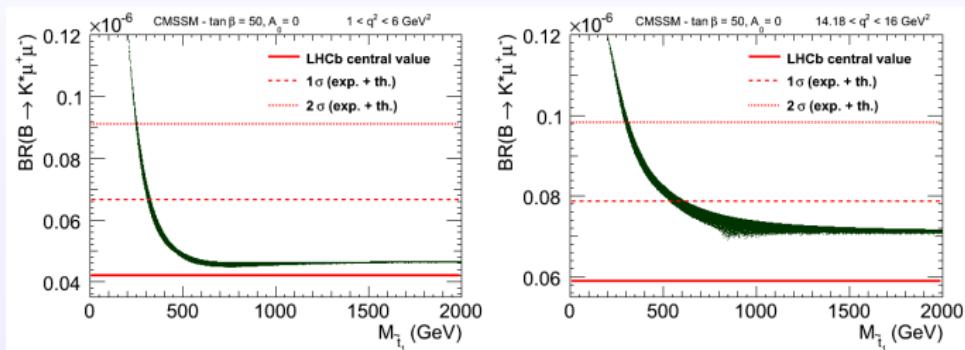
New LHCb limits for $BR(B_s \rightarrow \mu^+ \mu^-)$ and $BR(B_d \rightarrow \mu^+ \mu^-)$

SuperIso v3.3

Constraints on CMSSM – $B \rightarrow K^* \mu^+ \mu^-$

$\text{BR}(B \rightarrow K^* \mu^+ \mu^-)$ in the low and high q^2 regions:

CMSSM - $\tan \beta = 50$



FM, S. Neshatpour, J. Orloff, arXiv:1205.1845 [hep-ph]

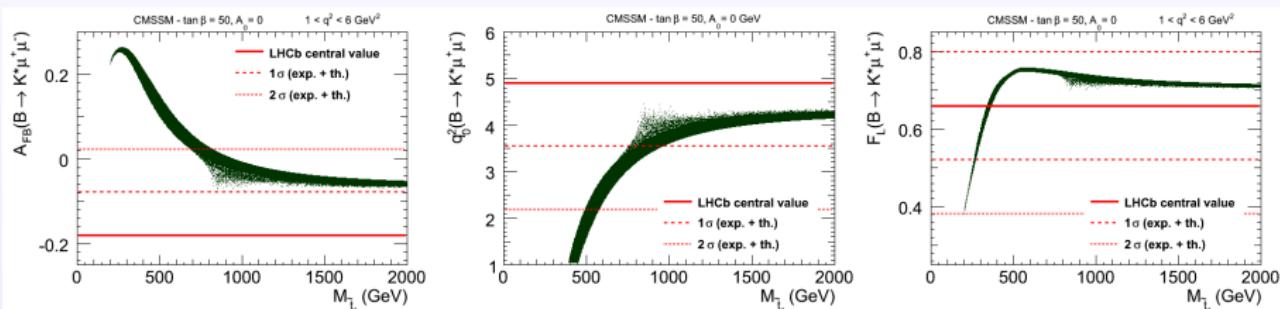
For $m_{\tilde{t}_1} > \sim 300 \text{ GeV}$, SUSY spread is within the th+exp error

- Look at other observables (A_{FB} , F_L , ...)
- Reduce both theory and experimental errors.

Constraints on CMSSM – $B \rightarrow K^* \mu^+ \mu^-$

Other observables of interest:

CMSSM - $\tan \beta = 50$



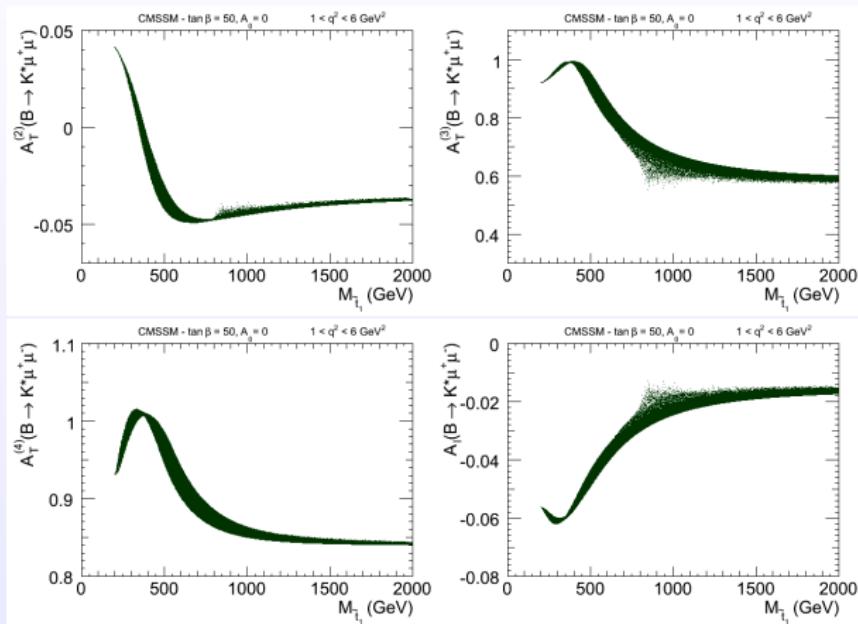
FM, S. Neshatpour, J. Orloff, arXiv:1205.1845 [hep-ph]

A_{FB} in the low q^2 region is especially interesting!

Constraints on CMSSM – $B \rightarrow K^* \mu^+ \mu^-$

Other observables (not yet measured):

CMSSM - $\tan \beta = 50$



FM, S. Neshatpour, J. Orloff, arXiv:1205.1845 [hep-ph]

Going beyond constrained scenarios

- CMSSM useful for benchmarking, model discrimination,...
- However the mass patterns could be more complicated

Phenomenological MSSM (pMSSM)

- The most general CP/R parity-conserving MSSM
- Minimal Flavour Violation at the TeV scale
- The first two sfermion generations are degenerate
- The three trilinear couplings are general for the 3 generations
 - 19 free parameters

10 sfermion masses, 3 gaugino masses, 3 trilinear couplings, 3 Higgs/Higgsino

A. Djouadi et al., hep-ph/9901246

→ Interplay between low energy observables and high p_T results

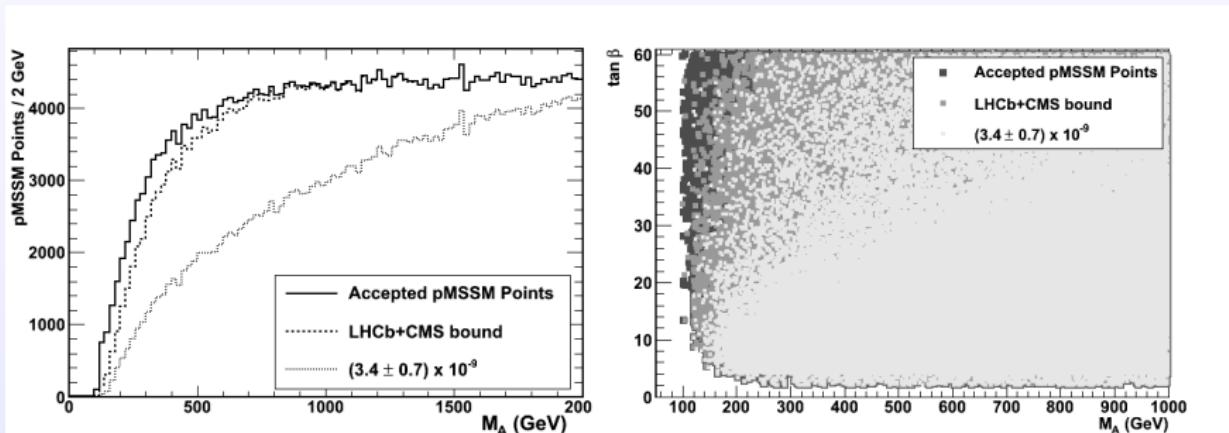
General MSSM – Sensitivity to M_A from $\text{BR}(B_s \rightarrow \mu^+ \mu^-)$

Considering 2 scenarios:

- 2011 bound from LHCb+CMS + estimated th syst:

$$\text{BR}(B_s \rightarrow \mu^+ \mu^-) < 1.26 \times 10^{-8}$$

- SM like branching ratio with estimated 20% total uncertainty



Light M_A strongly constrained!

- A. Arbey, M. Battaglia, F.M., Eur.Phys.J. C72 (2012) 1847
A. Arbey, M. Battaglia, F.M., Eur.Phys.J. C72 (2012) 1906

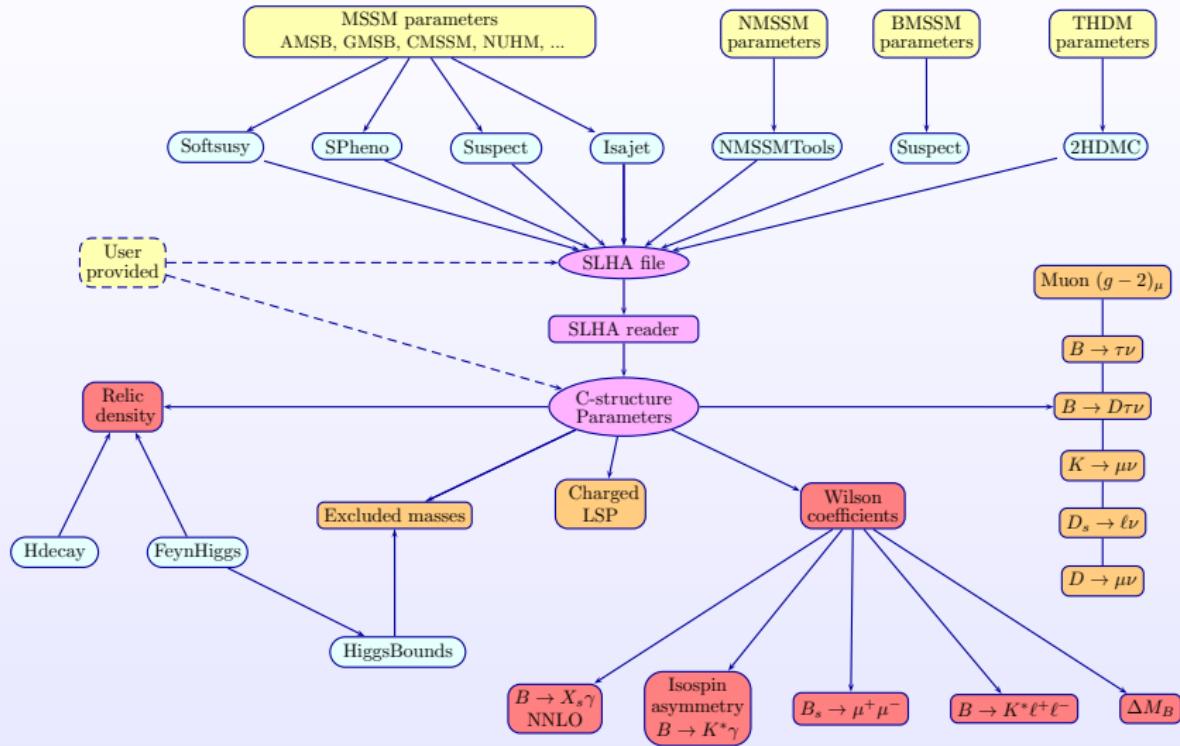
- public C program
- dedicated to the flavour physics observable calculations
- various models implemented
- interfaced to several spectrum calculators
- modular program with a well-defined structure
- complete reference manuals available

<http://superiso.in2p3.fr>

FM, Comput. Phys. Commun. 178 (2008) 745

FM, Comput. Phys. Commun. 180 (2009) 1579

FM, Comput. Phys. Commun. 180 (2009) 1718



- Flavour physics plays a very important role in constraining BSM scenarios
- $B_s \rightarrow \mu^+ \mu^-$ is a particularly sensitive to the scalar contributions and the high $\tan \beta$ regime
- $B \rightarrow K^* \mu^+ \mu^-$ offers multiple sensitive observables
→ complementary information!
- Theory uncertainties under control
- With more data constraints will tighten!

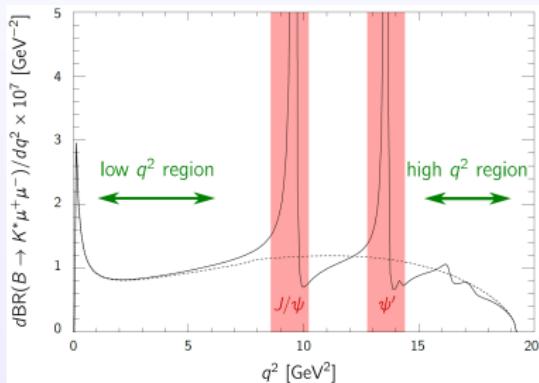
Backup

• Low q^2

- small $1/m_b$ corrections
- sensitivity to the interference of C_7 and C_9
- high rate
- long-distance effects not fully under control
- non-negligible scale and m_c dependence

• High q^2

- negligible scale and m_c dependence due to the strong sensitivity to C_{10}
- negligible long-distance effects of the type $B \rightarrow J/\Psi X_s \rightarrow X_s + X' \ell^+ \ell^-$
- sizable $1/m_b$ corrections
- low rate



Isospin asymmetry:

Non-factorizable graphs: annihilation or spectator-scattering diagrams

Isospin asymmetry arises when a photon is radiated from the spectator quark

→ depends on the charge of the spectator quark

→ different for charged and neutral B meson decays

$$\frac{dA_I}{dq^2} \equiv \frac{\frac{d\Gamma}{dq^2}(B^0 \rightarrow K^{*0} \ell^+ \ell^-) - \frac{d\Gamma}{dq^2}(B^- \rightarrow K^{*-} \ell^+ \ell^-)}{\frac{d\Gamma}{dq^2}(B^0 \rightarrow K^{*0} \ell^+ \ell^-) + \frac{d\Gamma}{dq^2}(B^- \rightarrow K^{*-} \ell^+ \ell^-)}$$

The SM is sensitive to C_5 and C_6 at small q^2 , but to C_3 and C_4 at larger q^2

Model independent constraints on New Physics: post-LHCb

Observable	Experiment	SM prediction
$\text{BR}(B \rightarrow X_s \gamma)$	$(3.55 \pm 0.24 \pm 0.09) \times 10^{-4}$	$(3.08 \pm 0.24) \times 10^{-4}$
$\Delta_0(B \rightarrow X_s \gamma)$	$(5.2 \pm 2.6 \pm 0.09) \times 10^{-2}$	$(8.0 \pm 3.9) \times 10^{-2}$
$\text{BR}(B \rightarrow X_d \gamma)$	$(1.41 \pm 0.57) \times 10^{-5}$	$(1.49 \pm 0.30) \times 10^{-5}$
$\text{BR}(B_s \rightarrow \mu^+ \mu^-)$	$< 4.5 \times 10^{-9}$	$(3.53 \pm 0.38) \times 10^{-9}$
$\langle d\text{BR}/dq^2(B \rightarrow K^* \mu^+ \mu^-) \rangle_{q^2 \in [1, 6] \text{ GeV}^2}$	$(0.42 \pm 0.04 \pm 0.04) \times 10^{-7}$	$(0.47 \pm 0.27) \times 10^{-7}$
$\langle d\text{BR}/dq^2(B \rightarrow K^* \mu^+ \mu^-) \rangle_{q^2 \in [14, 18, 16] \text{ GeV}^2}$	$(0.59 \pm 0.07 \pm 0.04) \times 10^{-7}$	$(0.71 \pm 0.18) \times 10^{-7}$
$\langle A_{FB}(B \rightarrow K^* \mu^+ \mu^-) \rangle_{q^2 \in [1, 6] \text{ GeV}^2}$	$-0.18 \pm 0.06 \pm 0.02$	-0.06 ± 0.05
$\langle A_{FB}(B \rightarrow K^* \mu^+ \mu^-) \rangle_{q^2 \in [14, 18, 16] \text{ GeV}^2}$	$0.49 \pm 0.06 \pm 0.05$	0.44 ± 0.10
$q_0^2(A_{FB}(B \rightarrow K^* \mu^+ \mu^-))$	$4.9^{+1.1}_{-1.3} \text{ GeV}^2$	$4.26 \pm 0.34 \text{ GeV}^2$
$\langle F_L(B \rightarrow K^* \mu^+ \mu^-) \rangle_{q^2 \in [1, 6] \text{ GeV}^2}$	$0.66 \pm 0.06 \pm 0.04$	0.71 ± 0.13
$\text{BR}(B \rightarrow X_s \mu^+ \mu^-)_{q^2 \in [1, 6] \text{ GeV}^2}$	$(1.60 \pm 0.68) \times 10^{-6}$	$(1.78 \pm 0.16) \times 10^{-6}$
$\text{BR}(B \rightarrow X_s \mu^+ \mu^-)_{q^2 > 14.4 \text{ GeV}^2}$	$(4.18 \pm 1.35) \times 10^{-7}$	$(2.18 \pm 0.65) \times 10^{-7}$

Model independent constraints on New Physics: pre-LHCb

Observable	Experiment
$\text{BR}(B \rightarrow X_s \gamma)$	$(3.55 \pm 0.24 \pm 0.09) \times 10^{-4}$
$\Delta_0(B \rightarrow X_s \gamma)$	$(5.2 \pm 2.6 \pm 0.09) \times 10^{-2}$
$\text{BR}(B \rightarrow X_d \gamma)$	$(1.41 \pm 0.57) \times 10^{-5}$
$\text{BR}(B_s \rightarrow \mu^+ \mu^-)$	$< 5.8 \times 10^{-8}$
$\langle d\text{BR}/dq^2(B \rightarrow K^* \ell^+ \ell^-) \rangle_{q^2 \in [1, 6] \text{ GeV}^2}$	$(0.32 \pm 0.11 \pm 0.03) \times 10^{-7}$
$\langle d\text{BR}/dq^2(B \rightarrow K^* \mu^+ \mu^-) \rangle_{q^2 \in [14, 18, 16] \text{ GeV}^2}$	$(0.83 \pm 0.20 \pm 0.07) \times 10^{-7}$
$\langle A_{FB}(B \rightarrow K^* \mu^+ \mu^-) \rangle_{q^2 \in [1, 6] \text{ GeV}^2}$	$0.43 \pm 0.36 \pm 0.06$
$\langle A_{FB}(B \rightarrow K^* \mu^+ \mu^-) \rangle_{q^2 \in [14, 18, 16] \text{ GeV}^2}$	$0.42 \pm 0.16 \pm 0.09$
$\langle F_L(B \rightarrow K^* \mu^+ \mu^-) \rangle_{q^2 \in [1, 6] \text{ GeV}^2}$	$0.50 \pm 0.30 \pm 0.03$
$\text{BR}(B \rightarrow X_s \mu^+ \mu^-)_{q^2 \in [1, 6] \text{ GeV}^2}$	$(1.60 \pm 0.68) \times 10^{-6}$
$\text{BR}(B \rightarrow X_s \mu^+ \mu^-)_{q^2 > 14.4 \text{ GeV}^2}$	$(4.18 \pm 1.35) \times 10^{-7}$

Double ratios of leptonic decays

$$R = \left(\frac{\text{BR}(B_s \rightarrow \mu^+ \mu^-)}{\text{BR}(B_u \rightarrow \tau \nu)} \right) / \left(\frac{\text{BR}(D_s \rightarrow \tau \nu)}{\text{BR}(D \rightarrow \mu \nu)} \right)$$

From the form factor and CKM matrix point of view:

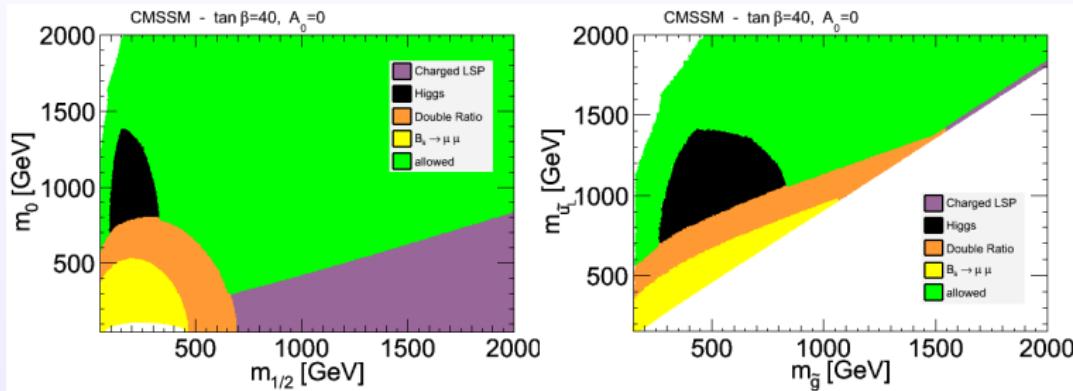
$$R \propto \frac{|V_{ts} V_{tb}|^2}{|V_{ub}|^2} \frac{(f_{B_s}/f_B)^2}{(f_{D_s}/f_D)^2} \quad \text{with:} \quad \frac{(f_{B_s}/f_B)}{(f_{D_s}/f_D)} \approx 1$$

R has no dependence on the decay constants, contrary to each decay taken individually!

- No dependence on lattice quantities
- Interesting for V_{ub} determination
- Interesting for probing new physics
- Promising experimental situation

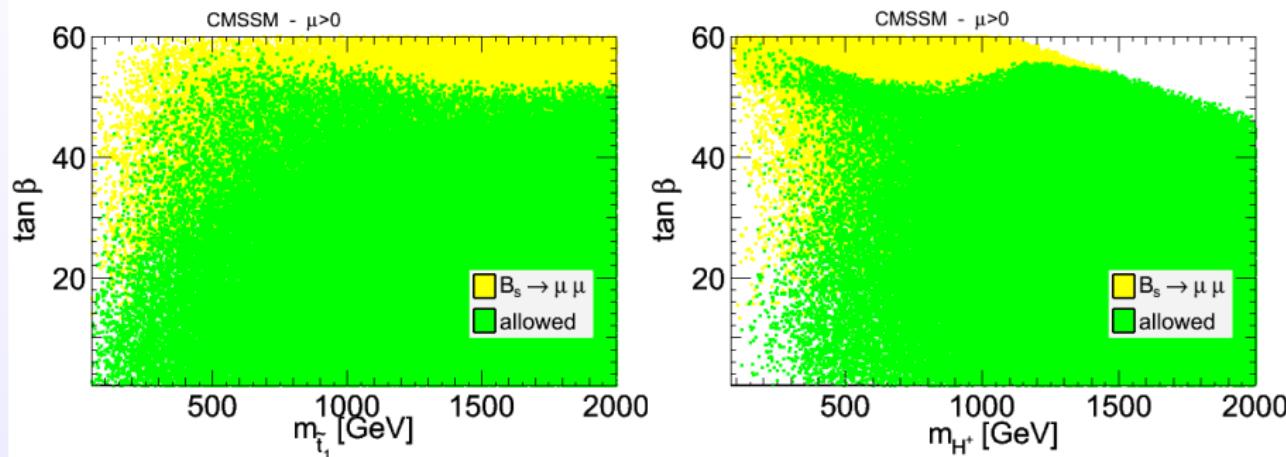
B. Grinstein, Phys. Rev. Lett. 71 (1993)
A.G. Akeroyd, FM, JHEP 1010 (2010)

Constrained MSSM



Implications on SUSY – $\text{BR}(B_s \rightarrow \mu^+ \mu^-)$

Constraints in CMSSM (all parameters varied)

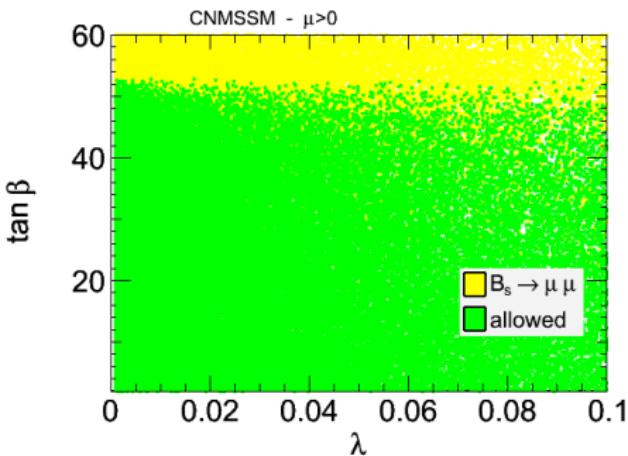
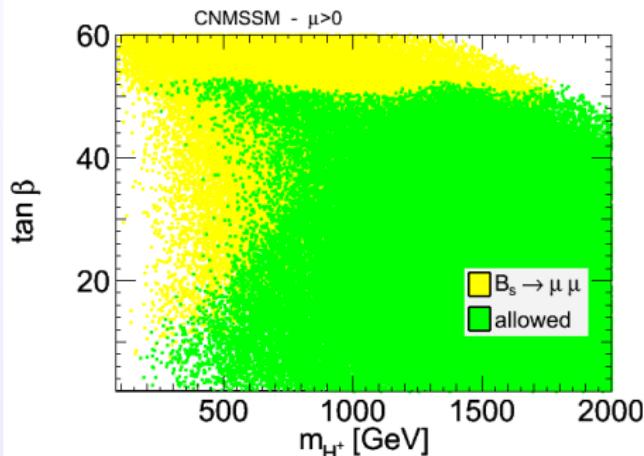


At 95% C.L., including th uncertainty: $\text{BR}(B_s \rightarrow \mu^+ \mu^-) < 5.0 \times 10^{-9}$

A.G. Akeroyd, F.M., D. Martinez Santos, JHEP 1112 (2011) 088
SuperIso v3.2

Implications on SUSY – $\text{BR}(B_s \rightarrow \mu^+ \mu^-)$

Constraints in CNMSSM (all parameters varied)



A.G. Akeroyd, F.M., D. Martinez Santos, JHEP 1112 (2011) 088
SuperIso v3.2

Muon anomalous magnetic moment

