Beyond Special Relativity and DSR theory

José Javier Relancio Martínez

Introduction

DSR model at first order in A

DSR model at second order in Λ and space-time

Conclusions

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Motivation

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Space-time: the last frontier

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Conclusions

- Any answer has to include matter and also the space-time structure \rightarrow Gravity
- If fundamental constituents of matter exist, does it happen the same for space-time?

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Do space «atoms» exist?

QFT y GR: incompatibilities

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Conclusions

- One of the challenges for the present physics is the unification of GR and QFT → QGT.
- In QFT, one asumes a given space-time and one studies in detail the properties and the movement of particles in it.

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 In GR, one asumes that the properties of matter and radiation are given and one describes in detail the resultant space-time (curvature).

QFT y GR: incompatibilities

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Why do we need a QGT?

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Study of the first moments of the universe.

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- Black holes: Singularity, information?
- Answers→ QGT.

Quantum Gravity Theories

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Conclusions

- Approaches to QGT: string theory, quantum loop gravity, supergravity, causal set theory...
- There are no experimental evidences of a fundamental QGT.
- New approach: study the low energy theory of QGT. Doubly Special Relativity (DSR) →posible experimental evidences.

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Doubly Special Relativity (DSR)

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Conclusions

- Two invariants in every inertial frame: speed of light c and Planck length I_P.
- We can obtain t_P , M_P y Λ

 $\frac{\Lambda}{c^2}$

$$l_P = \sqrt{rac{\hbar G}{c^3}} = 1, 6 imes 10^{-35} \,\mathrm{m}$$

 $t_P = \sqrt{rac{\hbar G}{c^5}} = 5, 4 imes 10^{-44} \,\mathrm{s}$
 $= M_P = \sqrt{rac{\hbar c}{G}} = 2, 2 imes 10^{-8} \,\mathrm{kg} = 1, 2 imes 10^{19} \,\mathrm{GeV}/c^2$

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Quantum Gravity Phenomenology

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Conclusions

- Planck energy $ightarrow 10^{19}$ GeV
- \blacksquare Particle accelerators \rightarrow 1.3 \times 10 4 GeV
- Cosmic rays $\rightarrow 10^{11}$ GeV.
- Phenomenology?→ Amplifications at low energy.

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Photon time delay

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Conclusions

- A «foamy» structure of space-time could produce stochastic variations of the speed of particles.
- These deviations can be obtained through modified dispersion relations (MDR), that for $E \ll \Lambda$

$$E^2 - \vec{p}^2 - m^2 \simeq \xi_n E^2 \left(\frac{E}{\Lambda}\right)^n$$

With the Hamiltonian concept of speed

$$v = \frac{dE}{dp}$$

this causes a difference in the flight time

$$\Delta t \sim \left(\frac{d}{c}\right) \xi_n \left(\frac{E}{\Lambda}\right)^n$$

Photon time delay

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- This delay can be measured for photons with different energies coming from a *gamma ray burst*.
- Recent experiments impose strong restrictions to deviations respect to SR at leading order (n = 1)



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Relative locality

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Conclusions

- When SR was formulated, space-time was flat.
- With the development of GR, Einstein introduces the concept of curved space-time.
- Born: why not to consider also a curved momentum space?
- This happens when one considers a modified composition law (MCL) of momenta

$$(p\oplus q)^\mu = p^\mu + q^\mu + \Gamma^\mu_{
u\lambda}p^
u q^\lambda + ...$$

where $\Gamma^{\mu}_{\nu\lambda}$ is the affine connection. This composition law is not neccessrely commutative

$$p\oplus q \neq q\oplus p$$

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Relative locality

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Conclusions

 Alice sees the production local and the detection non local, while Bob sees production non local and detection local.



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First order model

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Most general MDR rotational invariant at first order

$$p_0^2-ec{p}^2+rac{lpha_1}{\Lambda}p_0^3+rac{lpha_2}{\Lambda}ec{p}^2p_0=m^2$$

- Lorentz invariance violation (LIV) vs. M_P-Deformed Lorentz invariance (DSR)
- In DSR, the presence of an energy scale requires that the deformed transformations act non-linearly
- Then, a linear (p_0, \vec{p}) -conservation law is not compatible with a relativity principle (RP) \implies MCL.

$$p_{1} \oplus p_{2}|^{0} = p_{1}^{0} + p_{2}^{0} + \frac{\beta_{1}}{\Lambda} p_{1}^{0} p_{2}^{0} + \frac{\beta_{2}}{\Lambda} \vec{p}_{1} \vec{p}_{2}$$

$$\overrightarrow{p_{1} \oplus p_{2}} = \vec{p}_{1} + \vec{p}_{2} + \frac{\gamma_{1}}{\Lambda} p_{1}^{0} \vec{p}_{2} + \frac{\gamma_{2}}{\Lambda} \vec{p}_{1} p_{2}^{0} + \frac{\gamma_{3}}{\Lambda} \vec{p}_{1} \times \vec{p}_{2}$$

Lorentz Transformations

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Deformed infinitesimal Lorentz transformation (one particle sector):

 $\lambda_1\,,\,\lambda_2\,,\,\lambda_3$

Two particle sector

 $\eta_1^L, \eta_2^L, \eta_1^R, \eta_2^R$

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• Now we impose the Relativity Principle: all observers agree with conservation law.

Relations

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Conclusions

 One can obtain the relations between the coefficients of MDR and MCL and those of the Lorentz transformation.

$$\boldsymbol{\alpha_1} = -2\left(\lambda_1 + \lambda_2 + 2\lambda_3\right) \quad \boldsymbol{\alpha_2} = 2\left(\lambda_1 + 2\lambda_2 + 3\lambda_3\right)$$

$$\beta_1 = 2 (\lambda_1 + \lambda_2 + 2\lambda_3) \quad \beta_2 = -2\lambda_3 - \eta_1^L - \eta_1^R$$
$$\gamma_1 = \lambda_1 + 2\lambda_2 + 2\lambda_3 - \eta_1^L \quad \gamma_2 = \lambda_1 + 2\lambda_2 + 2\lambda_3 - \eta_1^R$$
$$\gamma_3 = \eta_2^L - \eta_2^R$$

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Relations

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Conclusions

From the previous relations, one can obtain the relations (golden rules) between the MDR and the MCL:

$$\alpha_1 = -\beta_1 \qquad \alpha_2 = \gamma_1 + \gamma_2 - \beta_2$$

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Not every coefficients satisfy a RP.

Modified Heisenberg algebra: Snyder algebra

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The Snyder algebra is

$$[x_{\mu}, x_{\nu}] = -\frac{i}{\Lambda^2} J_{\mu\nu} \qquad [x_{\mu}, p_{\nu}] = i \left(\eta_{\mu\nu} - \frac{1}{\Lambda^2} p_{\mu} p_{\nu} \right)$$

This algebra is related to momentum space with the following MCL (at leading order in Λ^2)

$$(p\oplus q)^{\mu}=p^{\mu}\!+\!q^{\mu}\!-\!rac{1}{\Lambda^{2}}\,(p.q)\,q^{\mu}\!-\!rac{1}{2\Lambda^{2}}\,(p.q)\,p^{\mu}\!-\!rac{1}{2\Lambda^{2}}p^{2}q^{\mu}$$

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Conclusions



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