

# Holography for $\mathcal{N} = 1^*$ on $S^4$

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- ▶ Apply gauge/gravity duality to this setup and test holography in a non-conformal setup.
- ▶ Study  $\mathcal{N} = 1$  theories holographically. Localization has not been successful for these theories.

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- ▶ Precision test of holography! In  $AdS_5/CFT_4$  one typically compares numbers. Here we have a whole function to match.
- ▶ Previous results for holography for  $\mathcal{N} = 1^*$  and  $\mathcal{N} = 2^*$  on  $\mathbb{R}^4$ . [FGPW], [GPPZ], [Pilch-Warner], [Buchel-Peet-Polchinski], [Evans-Johnson-Petrini], [Polchinski-Strassler], ... On  $S^4$  the holographic construction is more involved.

# Plan

- ▶  $\mathcal{N} = 1^*$  SYM theory on  $S^4$
- ▶ The gravity dual of  $\mathcal{N} = 2^*$  on  $S^4$
- ▶ Holographic calculation of  $F_{S^4}^{\mathcal{N}=2^*}$
- ▶ Holography for  $\mathcal{N} = 1^*$  SYM theory on  $S^4$
- ▶ Outlook

$\mathcal{N} = 1^*$  SYM theory on  $S^4$

$\mathcal{N} = 1^*$  SYM on  $\mathbb{R}^4$

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$$A_\mu, \quad X_{1,2,3,4,5,6}, \quad \lambda_{1,2,3,4}.$$

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Organize this into an  $\mathcal{N} = 1$  vector multiplet

$$A_\mu, \quad \psi_1 \equiv \lambda_4,$$

and 3 chiral multiplets

$$\chi_j = \lambda_j, \quad Z_j = \frac{1}{\sqrt{2}}(X_j + iX_{j+3}), \quad j = 1, 2, 3.$$



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The  $\mathcal{N} = 1^*$  theory is obtained by giving (independent) mass terms for the chiral multiplets.

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When there is a will there is a way! [Pestun], [Festuccia-Seiberg], ...

$$\begin{aligned}\mathcal{L}_{\mathcal{N}=1^*}^{S^4} &= \mathcal{L}_{\mathcal{N}=4}^{S^4} \\ &+ \frac{2}{R^2} \text{tr} (Z_1 \tilde{Z}_1 + Z_2 \tilde{Z}_2 + Z_3 \tilde{Z}_3) \\ &+ \text{tr} (m_1 \tilde{m}_1 Z_1 \tilde{Z}_1 + m_2 \tilde{m}_2 Z_2 \tilde{Z}_2 + m_3 \tilde{m}_3 Z_3 \tilde{Z}_3) \\ &- \frac{1}{2} \text{tr} (m_1 \chi_1 \chi_1 + m_2 \chi_2 \chi_2 + m_3 \chi_3 \chi_3 + \tilde{m}_1 \tilde{\chi}_1 \tilde{\chi}_1 + \tilde{m}_2 \tilde{\chi}_2 \tilde{\chi}_2 + \tilde{m}_3 \tilde{\chi}_3 \tilde{\chi}_3) \\ &- \frac{1}{\sqrt{2}} \text{tr} [m_i \epsilon^{ijk} Z_i \tilde{Z}_j \tilde{Z}_k + \tilde{m}_i \epsilon^{ijk} \tilde{Z}_i Z_j Z_k] \\ &+ \frac{i}{2R} \text{tr} (m_1 Z_1^2 + m_2 Z_2^2 + m_3 Z_3^2 + \tilde{m}_1 \tilde{Z}_1^2 + \tilde{m}_2 \tilde{Z}_2^2 + \tilde{m}_3 \tilde{Z}_3^2) .\end{aligned}$$

15 (real) relevant couplings in the Lagrangian + 1 complex gaugino vev + 1 complexified gauge coupling. A 19 (real) parameter family of  $\mathcal{N} = 1$  deformations of  $\mathcal{N} = 4$  SYM. Only 18 of these parameters are visible as modes in IIB supergravity.

For  $m_3 = \tilde{m}_3 = 0$ ,  $m_1 = m_2 \equiv m$  and  $\tilde{m}_1 = \tilde{m}_2 \equiv \tilde{m}$  we get the  $\mathcal{N} = 2^*$  theory.

## Results from localization for $\mathcal{N} = 2^*$

The  $SU(N)$  gauge symmetry is generally broken to  $U(1)^{N-1}$  by a vev for  $Z_3$ .

$$Z_3 = \text{diag}(a_1, \dots, a_N).$$

After supersymmetric localization the path integral for the theory on  $S^4$  reduces to a finite dimensional integral over the Coulomb moduli  $a_i$ . [Pestun]

$$\mathcal{Z} = \int \prod_{i=1}^N da_i \delta\left(\sum_{i=1}^N a_i\right) \prod_{i<j} (a_i - a_j)^2 \mathcal{Z}_{1\text{-loop}} |\mathcal{Z}_{\text{inst}}|^2 e^{-S_{\text{cl}}},$$

where

$$S_{\text{cl}} = \frac{8\pi^2 N}{\lambda} \sum_{i=1}^N a_i^2, \quad \mathcal{Z}_{1\text{-loop}} = \prod_{i=1}^N \frac{H^2(a_i - a_j)}{H(a_i - a_j + mR)H(a_i - a_j - mR)},$$

$$H(x) = \prod_{n=1}^{\infty} \left(1 + \frac{x^2}{n^2}\right)^n e^{-x^2/n}.$$

The function  $\mathcal{Z}_{\text{inst}}$  is Nekrasov's partition function with parameters  $\varepsilon_1 = \varepsilon_2 = 1/R$ .

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This answer depends on the regularization scheme. The scheme independent quantity is

$$\boxed{\frac{d^3 F_{S^4}}{d(mR)^3} = -2N^2 \frac{mR((mR)^2 + 3)}{((mR)^2 + 1)^2}}$$

This is the unambiguous result one can aim to compute holographically.

The gravity dual

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The result is that only 3 real scalars, 2 Abelian gauge fields and the metric survive in the bosonic sector. The 3 scalars,  $\{\phi, \psi, \chi\}$ , are dual to 3 relevant operators

$$\mathcal{O}_\phi, \quad \Delta_{\mathcal{O}_\phi} = 2; \quad \mathcal{O}_\psi, \quad \Delta_{\mathcal{O}_\psi} = 3; \quad \mathcal{O}_\chi, \quad \Delta_{\mathcal{O}_\chi} = 2.$$



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It is convenient to use

$$\eta = e^{\phi/\sqrt{6}}, \quad z = \frac{1}{\sqrt{2}}(\chi + i\psi), \quad \tilde{z} = \frac{1}{\sqrt{2}}(\chi - i\psi).$$

In Euclidean signature the fields  $z$  and  $\tilde{z}$  are independent.

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In Euclidean signature the fields  $z$  and  $\tilde{z}$  are independent.

For  $\chi = 0$  recover the truncation for  $\mathcal{N} = 2^*$  on  $\mathbb{R}^4$ . [Pilch-Warner]

## Supergravity setup

The Euclidean Lagrangian is

$$\mathcal{L} = \frac{1}{2\kappa^2} \left[ -\mathcal{R} + 12 \frac{\partial_\mu \eta \partial^\mu \eta}{\eta^2} + 4 \frac{\partial_\mu z \partial^\mu \tilde{z}}{(1 - z\tilde{z})^2} + \mathcal{V} \right],$$
$$\mathcal{V} \equiv -\frac{4}{L^2} \left( \frac{1}{\eta^4} + 2\eta^2 \frac{1 + z\tilde{z}}{1 - z\tilde{z}} + \frac{\eta^8}{4} \frac{(z - \tilde{z})^2}{(1 - z\tilde{z})^2} \right).$$

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To preserve the isometries of  $S^4$  take the “domain-wall” Ansatz

$$ds^2 = L^2 e^{2A(r)} ds_{S^4}^2 + dr^2,$$
$$\eta = \eta(r), \quad z = z(r), \quad \tilde{z} = \tilde{z}(r).$$

The masses of the scalars around the  $AdS_5$  vacuum are

$$m_\phi^2 L^2 = m_\chi^2 L^2 = -4, \quad m_\psi^2 L^2 = -3.$$

# The BPS equations

Plug the Ansatz in the supersymmetry variations of the 5d  $\mathcal{N} = 8$  theory and use the “conformal Killing spinors” on  $S^4$

$$\hat{\nabla}_\mu \zeta = \frac{1}{2} \gamma_5 \gamma_\mu \zeta ,$$

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to derive the BPS equations

$$z' = \frac{3\eta'(z\tilde{z} - 1) [2(z + \tilde{z}) + \eta^6(z - \tilde{z})]}{2\eta[\eta^6(\tilde{z}^2 - 1) + \tilde{z}^2 + 1]} ,$$

$$\tilde{z}' = \frac{3\eta'(z\tilde{z} - 1) [2(z + \tilde{z}) - \eta^6(z - \tilde{z})]}{2\eta[\eta^6(z^2 - 1) + z^2 + 1]} ,$$

$$(\eta')^2 = \frac{[\eta^6(z^2 - 1) + z^2 + 1] [\eta^6(\tilde{z}^2 - 1) + \tilde{z}^2 + 1]}{9L^2\eta^2(z\tilde{z} - 1)^2} ,$$

$$e^{2A} = \frac{(z\tilde{z} - 1)^2 [\eta^6(z^2 - 1) + z^2 + 1] [\eta^6(\tilde{z}^2 - 1) + \tilde{z}^2 + 1]}{\eta^8(z^2 - \tilde{z}^2)^2} .$$

## UV expansion

The (constant curvature) metric on  $\mathbb{H}^5$  is

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Solving the BPS equations iteratively, order by order in the asymptotic expansion as  $r \rightarrow \infty$ , we find (with  $L = 1$ )

$$e^{2A} = \frac{e^{2r}}{4} + \frac{1}{6}(\mu^2 - 3) + \dots ,$$

$$\eta = 1 + e^{-2r} \left[ \frac{2\mu^2}{3} r + \frac{\mu(\mu + v)}{3} \right] + \dots ,$$

$$\frac{1}{2}(z + \tilde{z}) = e^{-2r} \left[ 2\mu r + v \right] + \dots ,$$

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## UV expansion

The (constant curvature) metric on  $\mathbb{H}^5$  is

$$ds_5^2 = dr^2 + L^2 \sinh^2 \left( \frac{r}{L} \right) ds_{S^4}^2 .$$

Solving the BPS equations iteratively, order by order in the asymptotic expansion as  $r \rightarrow \infty$ , we find (with  $L = 1$ )

$$e^{2A} = \frac{e^{2r}}{4} + \frac{1}{6}(\mu^2 - 3) + \dots ,$$

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Here  $\mu$  and  $v$  are integration constants. Think of them as the “source” and “vev” for the operator  $\mathcal{O}_\chi$ . Compare to field theory to identify  $\mu = imR$ .

## IR expansion

At  $r = r_*$  the  $S^4$  shrinks smoothly to zero size. Solve the BPS equations close to  $r = r_*$ , and require that the solution is smooth to find

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$$e^{2A} = (r - r_*)^2 + \frac{7\eta_0^{12} + 20}{81\eta_0^4} (r - r_*)^4 + \dots,$$

$$\eta = \eta_0 - \left( \frac{\eta_0^{12} - 1}{27\eta_0^3} \right) (r - r_*)^2 \left[ 1 - \left( \frac{85 + 131\eta_0^{12}}{810\eta_0^4} \right) (r - r_*)^2 + \dots \right],$$

$$\frac{1}{2}(z + \tilde{z}) = \sqrt{\frac{\eta_0^6 - 1}{\eta_0^6 + 1}} \left[ \frac{\eta_0^6}{\eta_0^6 + 2} - \frac{2\eta_0^8(4\eta_0^6 + 5)}{15(\eta_0^6 + 2)^2} (r - r_*)^2 + \dots \right],$$

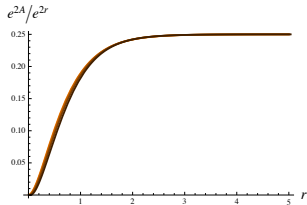
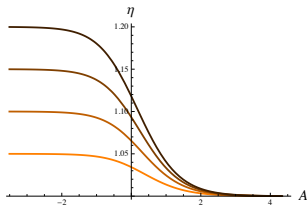
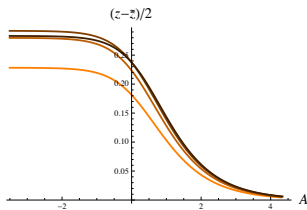
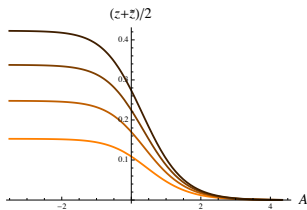
$$\frac{1}{2}(z - \tilde{z}) = \sqrt{\frac{\eta_0^6 - 1}{\eta_0^6 + 1}} \left[ \frac{2}{\eta_0^6 + 2} + \frac{\eta_0^2(3\eta_0^{12} - 10\eta_0^6 - 20)}{15(\eta_0^6 + 2)^2} (r - r_*)^2 + \dots \right].$$

Here,  $\eta_0$  is the only free parameter since one can set  $r_* = 0$  by the shift symmetry of the equations.

# Numerical solutions

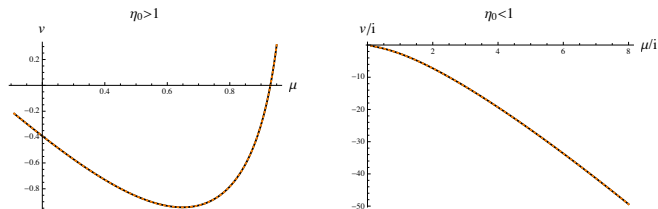
One can find numerical solutions by “shooting” from the IR to the UV. There is a one (complex) parameter family parametrized by  $\eta_0$ , so

$$v = v(\eta_0), \quad \text{and} \quad \mu = \mu(\eta_0).$$



# Numerical solutions

For real  $\eta_0$  one finds the following results for  $v(\mu)$



From the numerical results one can extract the following dependence

$$v(\mu) = -2\mu - \mu \log(1 - \mu^2)$$

Holographic calculation of  $F_{S^4}^{\mathcal{N}=2^*}$

# Calculating $F$ from supergravity

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- ▶ Without knowing this finite counterterm we can only hope to match  $\frac{d^5 F}{d\mu^5}$  with field theory.

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For both examples we found explicit supergravity truncations and (numerical) BPS solutions on  $S^4$ .

The hard part is to extract the third derivative of the free energy from the numerical solutions.

No results from localization to guide us.

## LS on $S^4$

From holographic renormalization one again finds

$$\frac{d^3 F^{\text{SUGRA}}}{d\mu^3} = -N^2 v''(\mu).$$

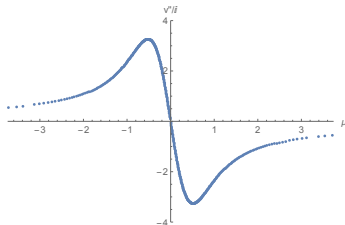
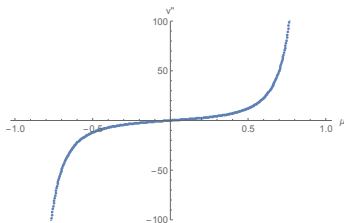
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On the whole complex  $\mu$ -plane  $v''(\mu)$  vanishes linearly around  $\mu \approx 0$ . It is an odd function of  $\mu$  and falls off for  $|\mu| \gg 1$  as

$$|v''(\mu)| \sim \frac{2}{|\mu|}.$$



There is a pole at  $\mu = \pm 1$  with

$$|v''(\mu)| \sim \frac{1}{(1 \pm \mu)^3}.$$

# Summary

- ▶ We found a 5d supegravity dual of  $\mathcal{N} = 2^*$  SYM on  $S^4$ .
- ▶ After careful holographic renormalization we computed the universal part of the free energy of this theory.
- ▶ The result is in exact agreement with the supersymmetric localization calculation in field theory.
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[Azeyanagi-Hanada-Honda-Matsuo-Shiba]

# Outlook

- ▶ Uplift of the  $\mathcal{N} = 2^*$  solution to IIB supergravity. Will allow for a holographic calculation of Wilson or 't Hooft line vevs. In addition one can study probe D3-branes. [Chen-Lin-Zarembo]
- ▶ Holography for  $\mathcal{N} = 1^*$  on other 4-manifolds. [Cassani-Martelli]
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- ▶ Extensions to other dimensions.
- ▶ Can we see some of the large N phase transitions argued to exist by Russo-Zarembo in IIB string theory?
- ▶ Revisit supersymmetric localization for  $\mathcal{N} = 1$  theories on  $S^4$ . Can one find the exact partition function (modulo ambiguities)? [Gerchkovitz-Gomis-Komargodski]
- ▶ For 4d  $\mathcal{N} = 2$  conformal theories  $Z_{S^4}$  gives the Zamolodchikov metric. What is the “meaning” of  $Z_{S^4}$  for non-conformal theories?
- ▶ Broader lessons for holography from localization?

**GRACIAS!**

## Comments

For a CFT on  $S^4$  of radius  $R$  with cutoff  $\epsilon \rightarrow 0$

$$F_{S^4} = \alpha_4 \left(\frac{R}{\epsilon}\right)^4 + \alpha_2 \left(\frac{R}{\epsilon}\right)^2 + \alpha_0 - a_{\text{anom}} \log\left(\frac{R}{\epsilon}\right) + \mathcal{O}(\epsilon/R).$$

For supersymmetric theories with a supersymmetric regularization scheme  $\alpha_4 = 0$ .

For massive theories there is an extra scale,  $m$ , so  $\alpha_2 = \alpha_2(m\epsilon)$  and  $\alpha_0 = \alpha_0(m\epsilon)$ . Expand this for small  $m\epsilon$

$$\begin{aligned}\alpha_2 &= \tilde{\alpha}_2 + m^2 \epsilon^2 \beta_2 + \mathcal{O}(m^4 \epsilon^4), \\ \alpha_0 &= \tilde{\alpha}_0 + \mathcal{O}(m^2 \epsilon^2).\end{aligned}$$

The nonuniversal contribution to the free energy is

$$\tilde{\alpha}_2 \left(\frac{R}{\epsilon}\right)^2 + \tilde{\alpha}_0 + \beta_2 (mR)^2.$$

Thus 3 derivatives w.r.t.  $mR$  eliminate the ambiguity.

For nonsupersymmetric theories the 5th derivative of  $F_{S^4}$  w.r.t.  $mR$  is universal.